Gravi-Reggeons and cosmic neutrino-nucleon interactions in theory with extra dimension

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SUMMARY

- Warped extra dimension (ED) with the small curvature
- AdS5 metric vs. flat metric with one compact ED
- Graviton production in exclusive double diffractive event
- Trans-Planckian scattering on the brane
- Neutrino-nucleon interactions at ultra-high energies
- Conclusions





Warped Extra Dimension with the Small Curvature

Background (AdS₅) metric (Randall & Sundrum, 1999)

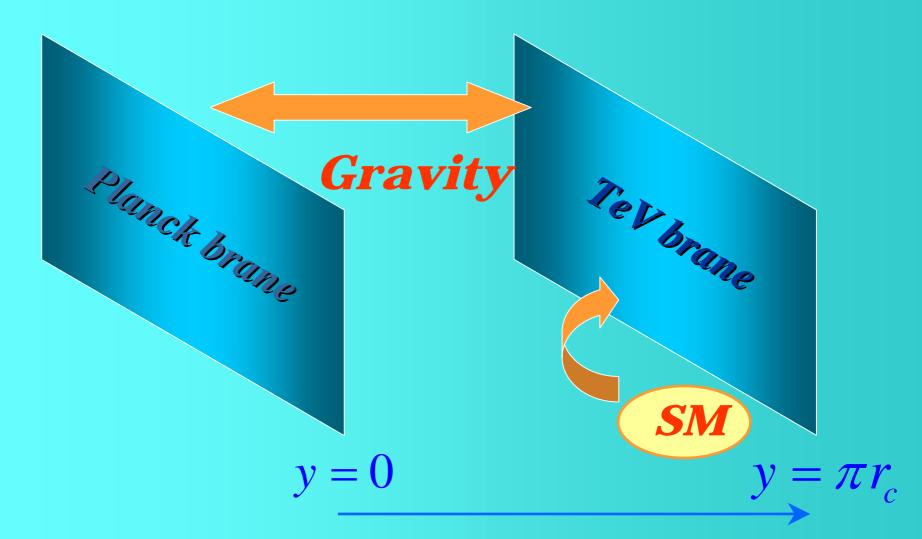
 $ds^{2} = \gamma_{MN}(z) dz^{M} dz^{N} = e^{2\kappa(|y| - \pi r)} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + dy^{2}$ $\eta_{\mu\nu} - Minkowski tensor \quad (M, N = 0, 1, ...4)$ $r \text{ is the radius of ED} \quad (-\pi r \le y \le \pi r)$

Hierarchy relation

$$M_{Pl}^2 = \frac{M_5^3}{\kappa} (e^{2\pi\kappa r} - 1) \quad \kappa r \approx 10$$

 M_{Pl} - Planck mass M_5 - gravity scale in 5 dimensions

Gravity lives in the bulk



SM fields are confined to the TeV brane

Gravitational 5-dimensional field

$$g_{MN}(z) = \gamma_{MN}(z) + \frac{2}{M_5^{3/2}} h_{MN}(z)$$

Kaluza-Klein (KK) gravitons
$$h_{\mu\nu}(x, y) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x) \psi_n(y)$$

Radion (scalar) field $h_{44}(x, y) = \phi(x)$

graviton masses
$$m_n = x_n \kappa$$

Interaction Lagrangian on the TeV brane

$$\mathbf{L} = -\left(\frac{1}{M_{Pl}}h_{\mu\nu}^{(0)} + \frac{1}{\Lambda_{\pi}}\sum_{n=1}^{\infty}h_{\mu\nu}^{(n)}\right)T^{\mu\nu} + \frac{1}{\sqrt{3}\Lambda_{\pi}}\phi T_{\alpha}^{\alpha}$$

physical scale
$$\Lambda_{\pi}^2 = \frac{M_5^3}{\kappa}$$

Large curvature option

 $\kappa \approx M_5 \approx 1 \,\mathrm{TeV}$

series of massive resonances

 $(\Lambda_{\pi}, m_1 \approx 1 \,\mathrm{TeV})$

Small curvature option
 $\kappa << M_5 \approx 1 \, \text{TeV}$ (Giudice et al., 2004,
Kisselev & Petrov, 2005)

narrow low-mass resonances with the small mass splitting ($\Delta m \approx \pi \kappa$)

$$10^{-5} \le \frac{\kappa}{M_5} \le 0.1$$

Formal relation to gravity in flat space-time with one compact ED

$$\kappa \rightarrow \frac{1}{\pi R_c}, \Lambda_{\pi} \rightarrow M_{Pl}$$
 R_c is the radius of

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AdS₅ Metric vs. Flat Metric with One Compact ED

RS model with the small curvature is not similar to a model with **one** large ED of the size $R_c^{-1} = \pi \kappa$

> For instance, $R_c^{-1} = 50 \text{ MeV}(1 \text{ GeV})$ can be realized only for d = 7(10)

> > d is the number of ED's

 $d = 1 \longrightarrow R_c = solar distance$

d = 2(3) - strongly limited by astrophysical bounds

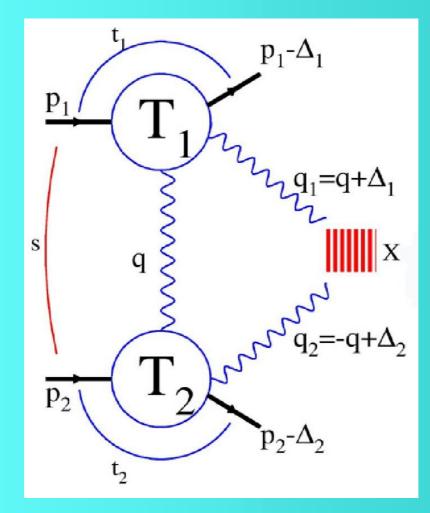
Hierarchy equation in d flat ED's (D=4+d)

$$M_{Pl}^{2} = \left(2\pi R_{c}\right)^{d} M_{D}^{2+d}$$

Limiting case of a similar relation for the warped metric

$$M_{Pl}^{2} = \frac{M_{5}^{3}}{\kappa} (e^{2\pi\kappa r} - 1) \xrightarrow{2\pi\kappa r \ll 1} 2\pi r M_{5}^{3}$$
$$2\pi\kappa r \ll 1 \longrightarrow \frac{M_{5}}{\kappa} \gg \left(\frac{M_{Pl}}{M_{5}}\right)^{2}$$
$$(unrealistic)$$

Graviton production in exclusive double diffractive event (EDDE)



X = KK graviton

Experimental signature at the LHC (Kisselev, Petrov & Ryutin, 2005)

$$p + p \rightarrow p + "nothing" + p$$

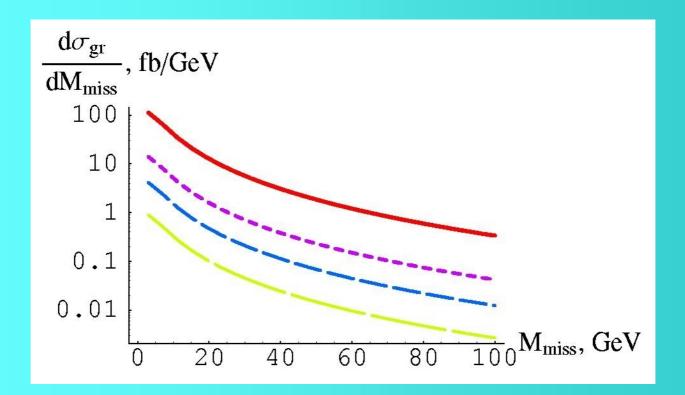
(joint CMS/TOTEM experiment)

Expected number of events

Integrated luminosity: $L = 30(0.3) \text{ fb}^{-1}$

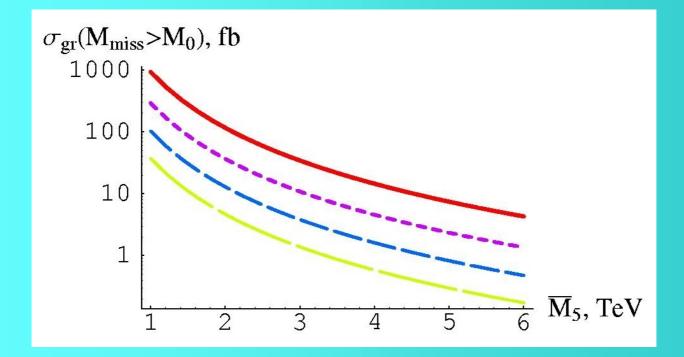
 $N_{ev} = 9000(90)$ for $M_5 = 1 \text{ TeV}$ and $M_{miss} > 14 \text{ GeV}$

Distribution in the missing mass of KK gravitons in EDDE



curves (top to bottom): $M_5 = 1, 2, 3 \text{ and } 5 \text{ TeV}$

Cross sections for the production of KK gravitons in EDDE



curves (top to bottom): $M_0 = 3, 14, 30$ and 50 GeV

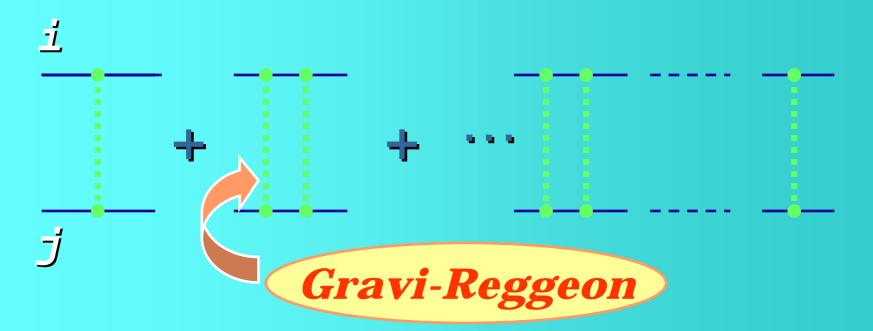
Trans-Planckian Scattering on the Brane

(Kisselev & Petrov, 2004/5)

Kinematical region **eikonal approximation** $\sqrt{s} \ge M_5, \quad \sqrt{s} >> -t$ *t* is 4-momentum transfer Born amplitude is the sum of the reggeized gravitons in t-channel

Gravi-Reggeons: $\alpha_n(t) = 2 + \alpha'_g t - \alpha'_g m_n^2$ $\alpha'_g = M_s^{-2}$ - string tension

Gravitational amplitude



<u> ゴ</u> , ブ - SM fields (q, g, l, n...)



Imaginary part of the eikonal

One has to calculate the sum

$$\sum_{n \neq 0} \exp\left(-\alpha'_{g}m_{n}^{2}\ln s\right)$$
with $m_{n} \approx \pi\kappa\left(n + \frac{1}{4}\right)$

At K ≤≤ M₃, we obtain

$$\operatorname{Im} \chi(s,b) \approx \frac{s}{M_5^2} \exp\left[-b^2 / 4R_g^2(s)\right]$$

 $R_g(s) = \sqrt{\alpha'_g \ln s}$ - gravitational radius

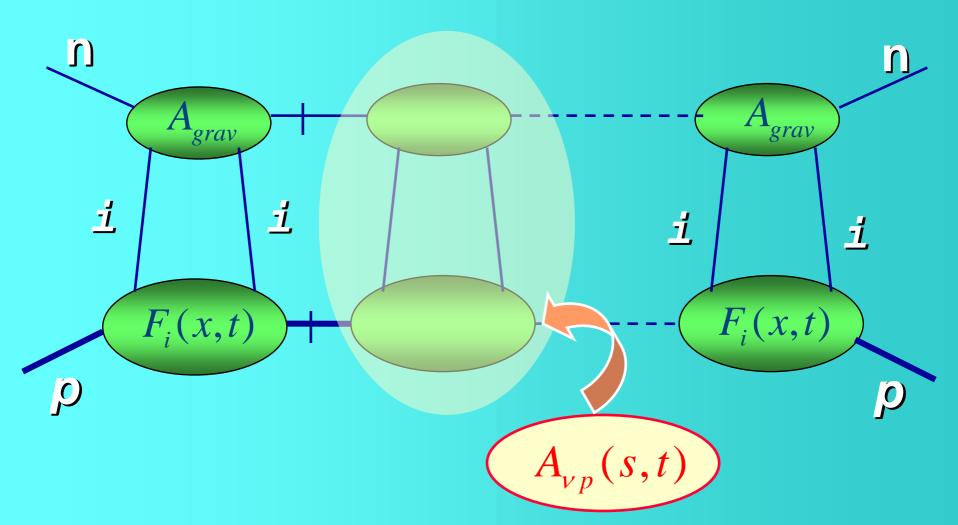
Fields are weakly coupled to gravity: $\Lambda_{\pi} \approx 100 \left(M_{5} / 1 \, TeV \right)^{3/2} \, \text{TeV}$ for small $\kappa = 100 \, \text{MeV}$

The summation of t-channel reggeized gravitons

eikonal with no explicit dependence on the brane scale and curvature

Radion production is strongly suppressed $\left(\Lambda_{\phi} = \sqrt{3}\Lambda_{\pi}\right)$

Neutrino-nucleon Interactions at Ultra-high Energies



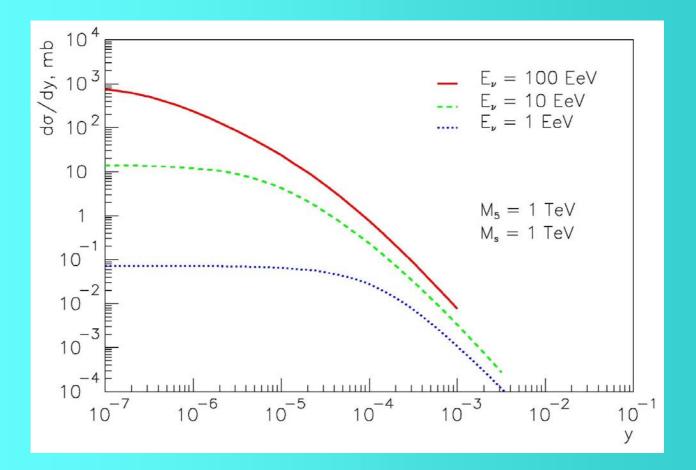
Gravitational amplitude
$$(\hat{s} = sx)$$

 $A_{grav}(\hat{s},t) \approx \frac{\hat{s}^2 \alpha'_g}{M_5^3 R_g(\hat{s})} \exp[t R_g(\hat{s})]$

Skewed (t-dependent) parton distribution $F_{i}(x,t) = f_{i}(x) \exp\left[t\left(b_{0}^{2} - \alpha_{P}' \ln x\right)\right]$ $f_{i}(x) - standard distribution$ of parton i

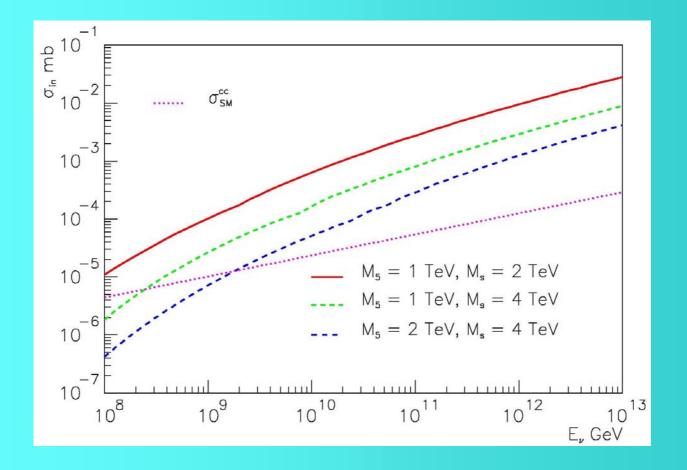
parameters of the hard Pomeron is used (Petrov & Prokudin, 2002)

Differential cross section



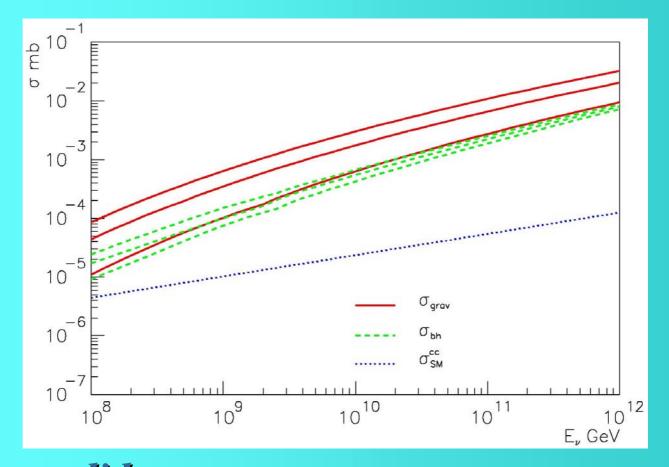
y - neutrino energy fraction deposited to the proton

notorq-onirtuen ottaslenl cross sections



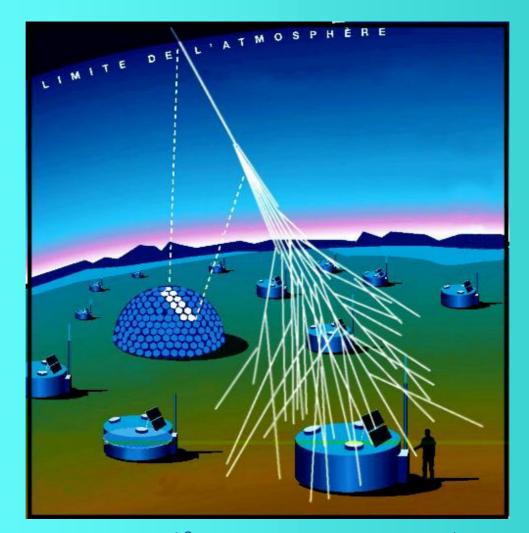
dotted curve: SM (cc) cross section

Neutrino-proton interaction vs. black hole production



solid curves: $M_5 = 0.25, 0.5, 1 \text{ TeV}$ dashed curves: $M_{BH}^{\min} = 0.5, 1, 2 \text{ TeV}$

Detection of quasi-horizontal air showers by the Pierre Auger Observatory



(from www.auger.org)

CONCLUSIONS

- Gravity effects from ED's may be detected in double diffractive events by the joint experiment of the CMS and TOTEM Collaborations at the LHC
- Trans-Planckian gravity induced scattering of the brane fields is given by an infinite sum of the t-channel reggeized gravitons (gravi-Reggeons)
- Gravitational part of the cross section for cosmic neutrino scattering off the nucleon is comparable with the BH production cross section

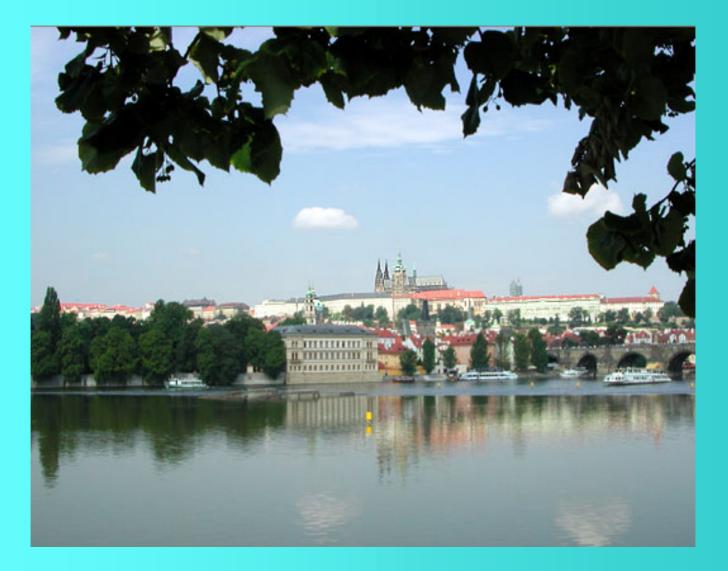
 Gravity effects from ED's are large enough to be measured in ultra-high-energy neutrino-nucleon events by presently planned neutrino detectors

Expected rate for quasi-horizontal air showers at the Auger Observatory:

4.9 per year $M_5 = 1 \text{ TeV}$
1.6 per year $M_5 = 2 \text{ TeV}$

(for Waxman-Bahcall neutrino flux)

SM: 0.08 per year



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