NEMO – toward a km³ Neutrino Telescope in the Mediterranean Sea

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Prague, Czech Republic From Colliders to Cosmic Rays 7-13 September 2005

OUTLINE

Introduction

The NEMO project
Description
Results
Perspectives

Conclusions

Physics Motivations

High Energy Neutrinos are a powerful tool to investigate the hadronic mechanisms of Cosmic Ray production

Galactic candidates:

- Pulsars
- SuperNova remants
- Micro-Quasars

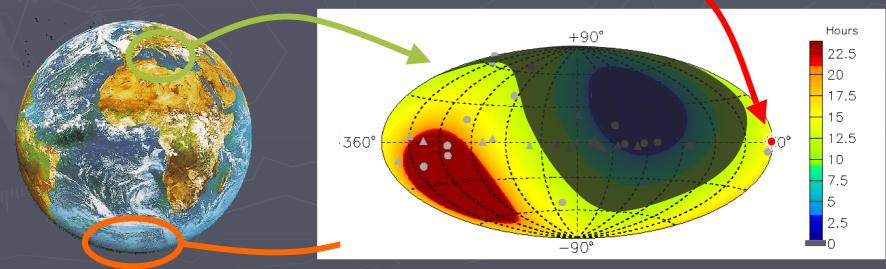
Extra-galactic candidates:

- Active Galactic Nuclei
- Gamma Ray Bursts

Physics Motivations

A neutrino telescope in the Northern Hemisphere will provide:

- Complementary sky survey to ICECUBE
- Overlap observation region with ICECUBE
- Study of the Galactic Center



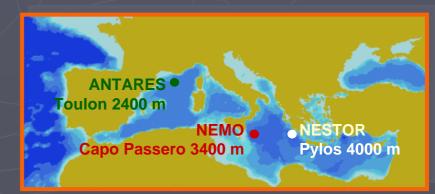
Physics Motivation

The small cross section and the expected low neutrino flux require

- Iarge volume telescope ~ 1 km³
- Iong observation time ~ yrs

The atmospheric muon background requires
 a shielding > 2000 m water equivalent

The Mediterranean Sea provides convenient sites



The NEMO Collaboration

FN INFN

- Bari, Bologna, Catania, Genova, LNF, LNS, Napoli, Pisa, Roma
- Universities
 - Bari, Bologna, Catania, Genova, Napoli, Pisa, Roma "La Sapienza"

CNR

- Istituto di Oceanografia Fisica, La Spezia
- Istituto di Biologia del Mare, Venezia
- Istituto Sperimentale Talassografico, Messina



Istituto Nazionale di Geofisica e Vulcanologia (INGV)



Istituto Nazionale di Oceanografia e Geofisica Sperimentale (OGS)



Istituto Superiore delle Comunicazioni e delle Tecnologie dell'Informazione (ISCTI)

more than 70 researchers involved

Technical Requirements

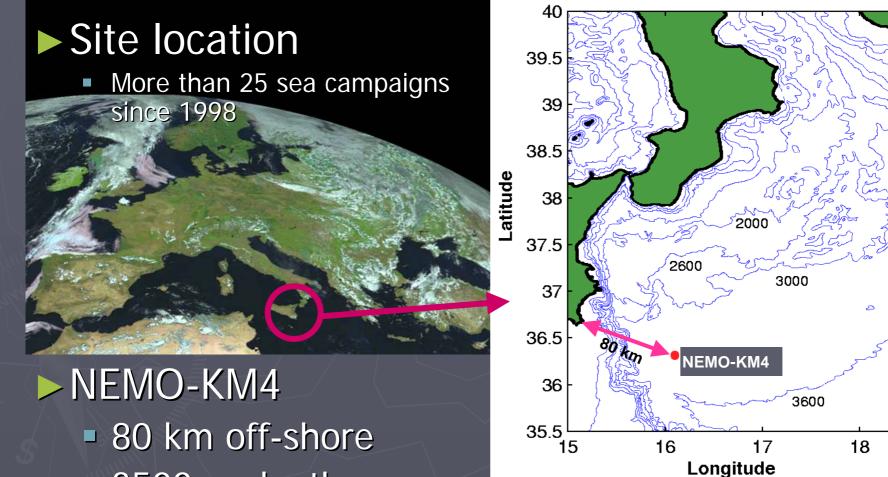
Issues to be solved to realize a km³ neutrino telescope in the Mediterranean Sea

- Best site location (i.e. depth, water quality ...)
- Optical background from ⁴⁰K-decay and bioluminescence
- Detector deployment
- Detector rigidity against sea currents
- Resistance to corrosion from salt water

The NEMO Project

- R&D phase (1999-2002)
 - Site selection and characterization
 - Several sites close to the italian coasts have been studied.
 - R&D Activities
 - Development of dedicated ASICS for the underwater front-end electronics
 - Development of large area hybrid photomultipliers
 - Feasibility Studies
 - All detector critical components and the deployment procedures have been examined
 - A preliminary project for the km³ detector has been developed
- Phase-1 and prototyping (2002-2006)
 - Realization and deployment of a prototype including all critical components
- Phase-2 (2006-...)
 - Realization of an underwater infrastructure at -3500 m

NEMO R&D Activity



3500 m depth

NEMO-KM4 Properties

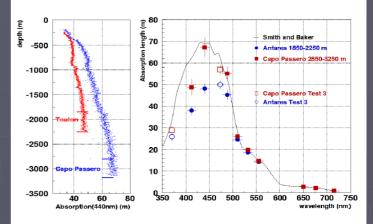
High water transparency

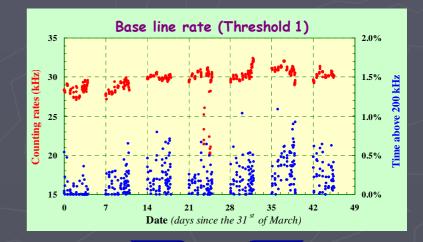
- Data compatible with pure salt water properties
- No seasonal variations

Reduced background, mostly from ⁴⁰K decay

- 10" PMT thres. 0.5 p.e. noise rate ~ 30 kHz
- Bioluminescence almost absent

Geologically stable





NEMO km³ Conceptual Design

electro optical cable: construction and deployment

Electronics

Power Distribution

Underwater connections

Power transmission system

Detector: design and construction deployment and recovery

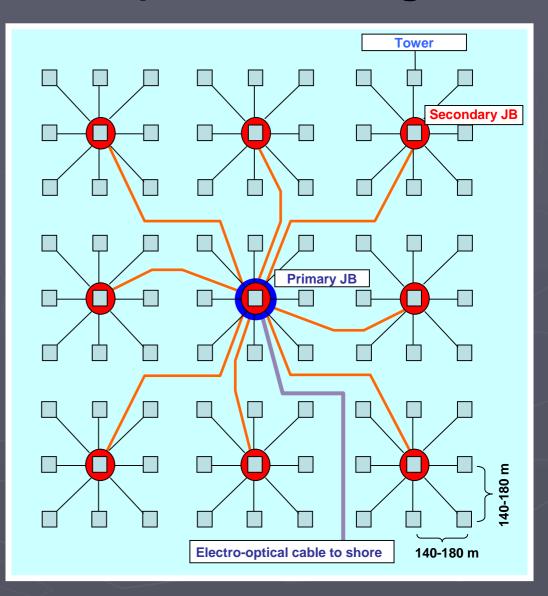
Data transmission system

Acoustic positioning

NEMO km³ Conceptual Design

Proposed lay-out

- 10 junction boxes
- 81 towers
- 5832 PMTs



Expected Performance

Simulations show excellent angular resolution and sensitivity

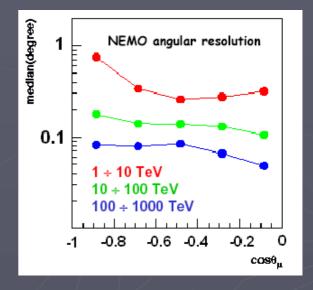
NEMO

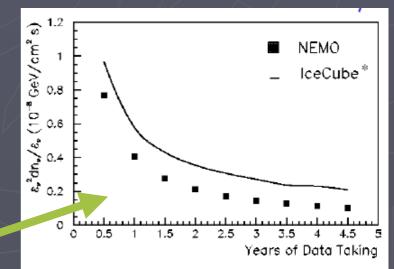
81 towers 140 m spaced5832 PMTs

ICECUBE

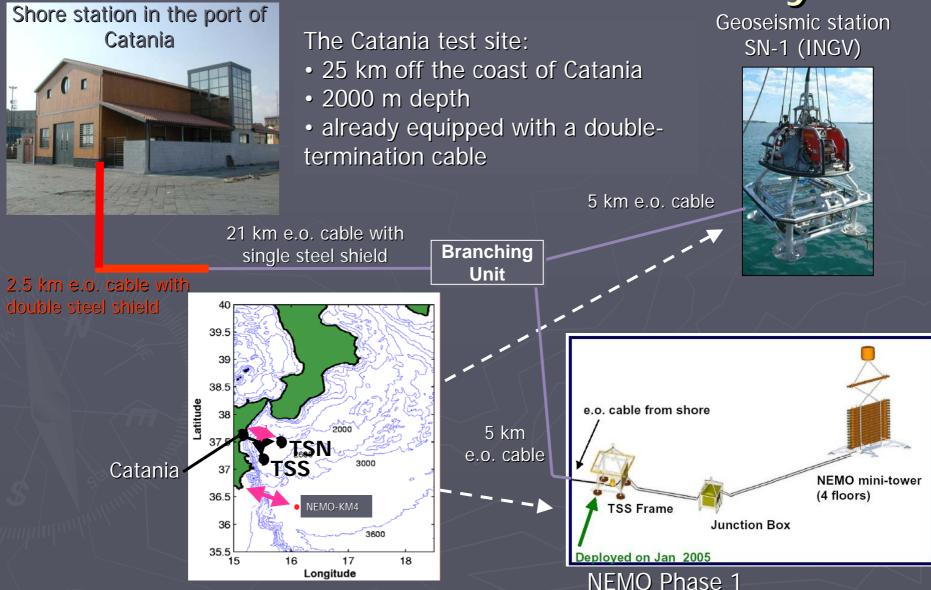
- 80 strings 125 m spaced
- ▶ 4800 PMTs

Sensitivity to a E⁻² neutrino spectrum from a pointlike source





The NEMO Phase-1 Activity



The NEMO Phase-1 Activity

Installation of the cable termination frames with electrooptical connectors



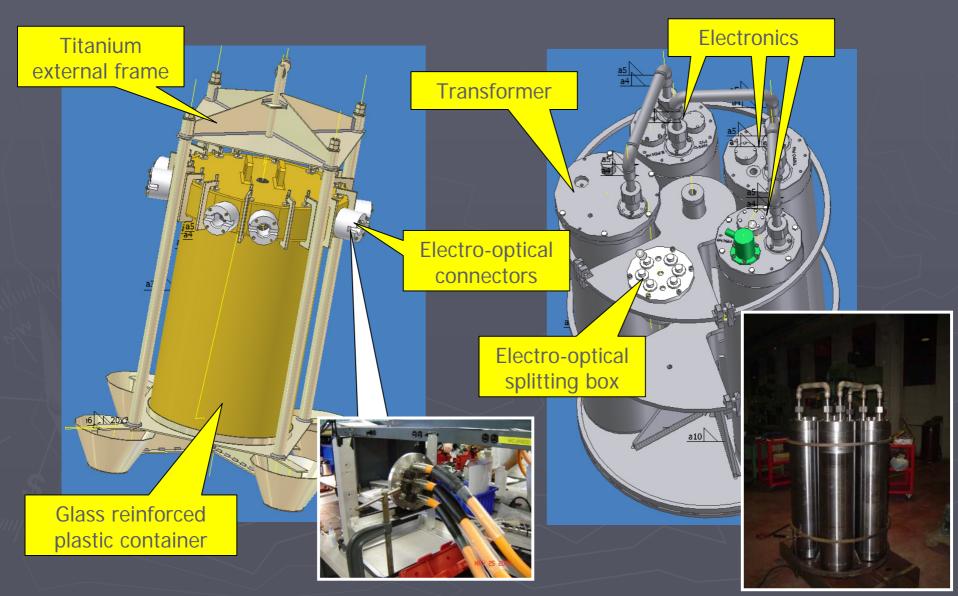
Deployment and connection of

- Acoustic detection station
- INGV environmental observatory

Fully operational since January 2005

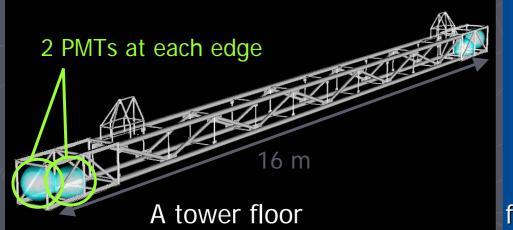


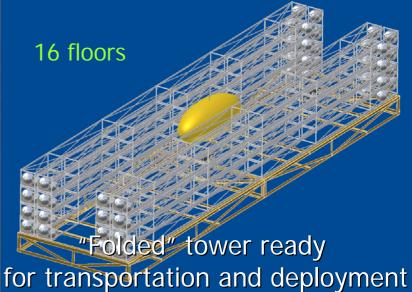
The NEMO Junction Box



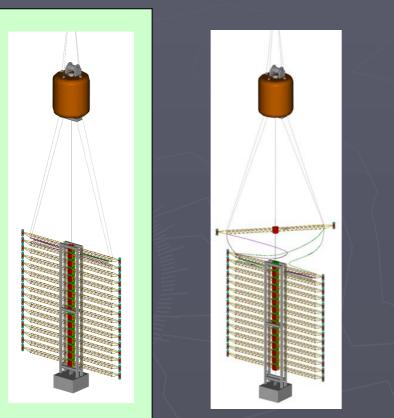
The NEMO Tower

 Semi-rigid structure provides "easy" assembly, transportation and deployment
 A 1:5 4-floor prototype has been successfully deployed and recovered in Spring 2004

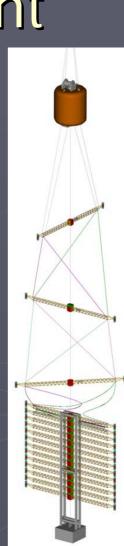


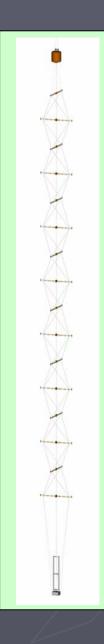


The Tower Deployment







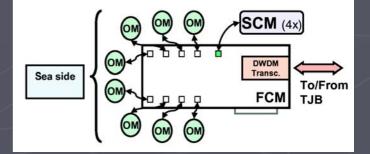


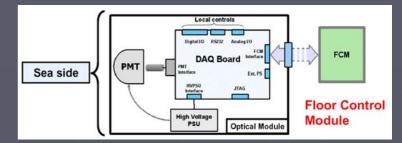
The Read-out Electronics

Optical Module Electronics
 200 Msample/s
 8 bit range (log. compression)

Floor Control Module Electronics Collects digitized signals from up to 8 OM Sends to the onshore station through the Junction Box









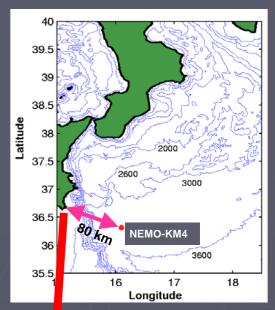
The NEMO Phase-2

Infrastructures for an -3500 m underwater site

- Electro-optical cable (40 kW)
 Purchase under way
- Shore station in PortoPalo di Capo Passero

Building under renovation







to NEMO-KM4

Conclusions and Perspectives

- The NEMO Collaboration is working on a long-term R&D program toward a km³ v-telescope in the Mediterranean Sea
 - An optimal candidate site has been found: NEMO-KM4
 - The NEMO Phase-1, aiming to validate the proposed technologies, is under way at the Catania Test Site
 - Since Jan. 2005 the geoseismic and acoustic stations are fully operative
 - The completion is planned in the first half of 2006
 - The NEMO Phase-2, aiming to realize the deep sea station at NEMO-KM4, is in progress
 - ► The purchase of the electro-optical cable is in progress
 - The set-up of the onshore station is in progress
 - ► The deployment of a full-size tower is foreseen in 2007

Toward the v-Telescope

EU is funding the joint activity for an Europeanscale Design Study for a km³ v-telescope in the Mediterranean Sea

KM3NeT: ANTARES-NEMO-NESTOR consortium

2° VLVnT (Very Large Volume v-Telescope) Workshop to be held in Catania (Italy) 8-11 Nov., 2005