

Understanding the emission duration through femtoscopy and imaging

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Workshop on Particle
Correlations and
Femtoscopia
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Talk Outline:

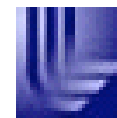
- Background
- Simple model emission function
- Source functions in this model
- Can we see this in experiment?
- CorAL Status



This work was performed under the auspices of the U.S.
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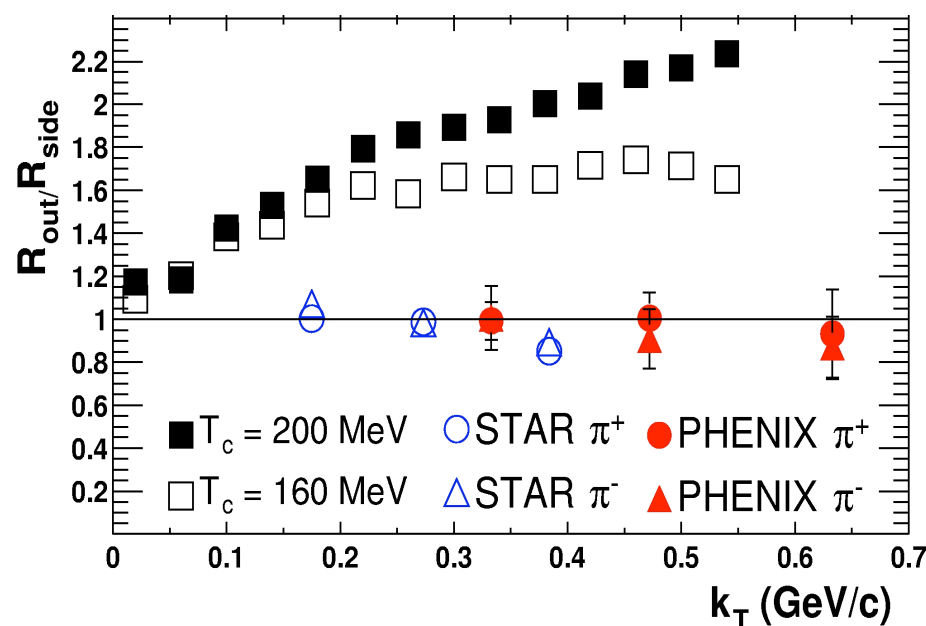
UCRL-PRES-214483

Mike's HBT Puzzle #1



What does QGP do to HBT radii? (Rischke, Gyulassy Nucl. Phys. A **608**, 479 (1996), many others):

- If there is a phase-transition, hydro evolution will slow in mixed phase.
- Will lead to long-lived source
- Huge difference in Outward/Sideward radii



$R_{out}/R_{side} == 1!$ (in LCMS)
Experiment analysis wrong?

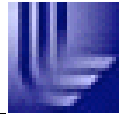
- Coulomb correction?

Theory screwed up?

Interpretation confused?

- Instant freeze-out?
- Gaussian fits?

The Koonin-Pratt Formalism



$$C(\vec{q}) = \int d^3r |\Phi_{\vec{q}}(\vec{r})|^2 S(\vec{r})$$

The pair's final state wave-function, we can compute this if we know the interaction. This is the kernel $K(q,r)$.

The *source function*, it gives the probability of producing a pair a distance r apart in the pair CM frame

We invert to get $S(r)$ directly

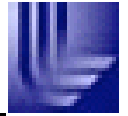
Source function related to emission function:

$$S_{\vec{P}}(\vec{r}') = \int dr'_0 \int d^4R D(R + r/2, \vec{P}/2) D(R - r/2, \vec{P}/2)$$

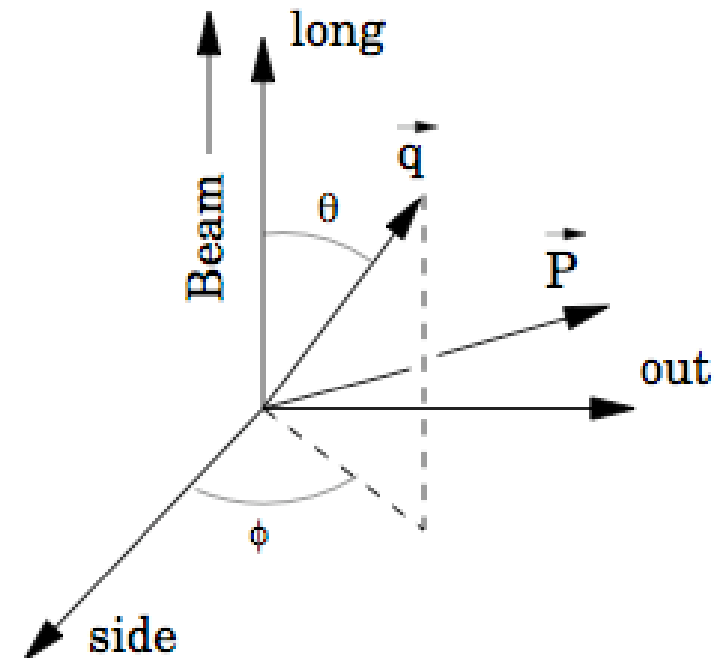
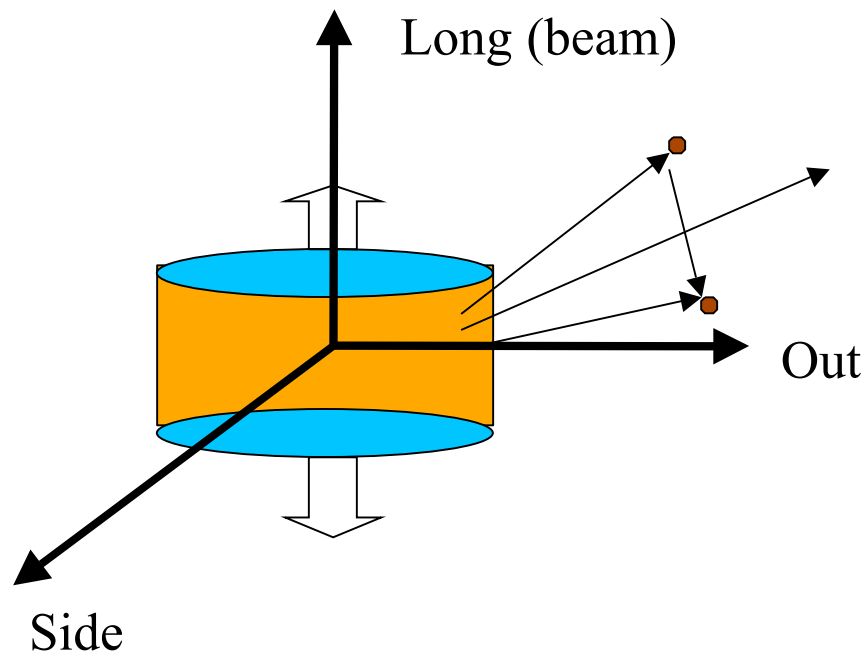


We work in Bertsch-Pratt coords., in pCM

Our coordinate choice

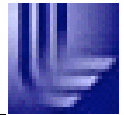


Work in Bertsch-Pratt coordinates in pair CM frame:

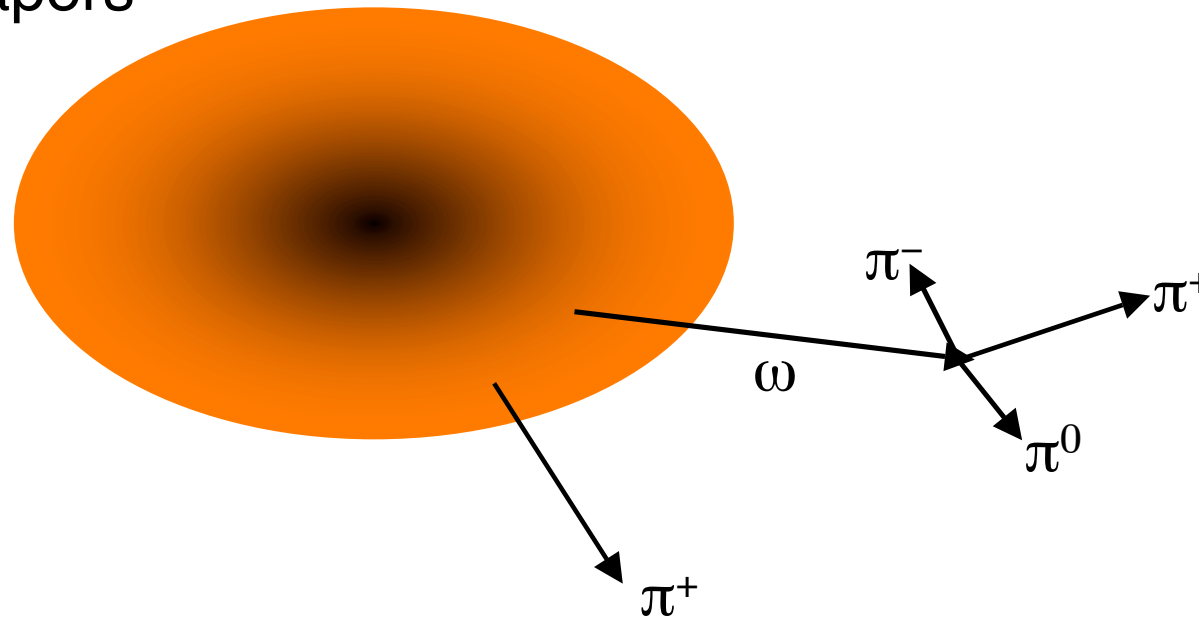


Boost from lab \rightarrow pair CM means lifetime effects transformed into Outwards/Longitudinal direction.

Core Halo Model



Nickerson, Csörög, Kiang, Phys. Rev. C **57**, 3251 (1998)
+ other papers

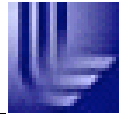


$$D(\mathbf{r}, t, \mathbf{p}) = f D_{\text{core}}(\mathbf{r}, t, \mathbf{p}) + (1 - f) D_{\text{halo}}(\mathbf{r}, t, \mathbf{p})$$

f is fraction of π 's emitted directly from core,
effectively $\lambda = f^2$ in source function



Core Halo Model, cont.



- From exploding core, with Gaussian shape:

$$D_{\text{core}}(\mathbf{r}, t, \mathbf{p}) \propto e^{-(E - \mathbf{p} \cdot \mathbf{v}_{\text{flow}})/T} D_{\text{gauss}}(\mathbf{r}) e^{-t/\tau_{fo}}$$

- Flow profile simple, but adjustable:

$$\mathbf{v}_{\text{flow}} = \begin{cases} \alpha \mathbf{r} & \text{if } |\mathbf{r}| < R_{\text{Au}} = 6.98 \text{ fm} \\ \frac{2}{3} c \hat{\mathbf{r}} & \text{otherwise} \end{cases}$$

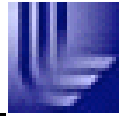
- From decay of emitted resonances, ω has most likely lifetime (23 fm/c)

$$D_{\text{halo}}(\mathbf{r}, t, \mathbf{p}) \propto \int d\Delta t d^3p_{\omega} P(\mathbf{p}_{\omega}, \mathbf{p}) e^{-\Delta t/\tau_{\omega}} \\ \times D_{\text{core}}\left(\mathbf{r} - \frac{\mathbf{p}_{\omega}}{E_{\omega}} \Delta t, t - \Delta t, \mathbf{p}_{\omega}\right)$$

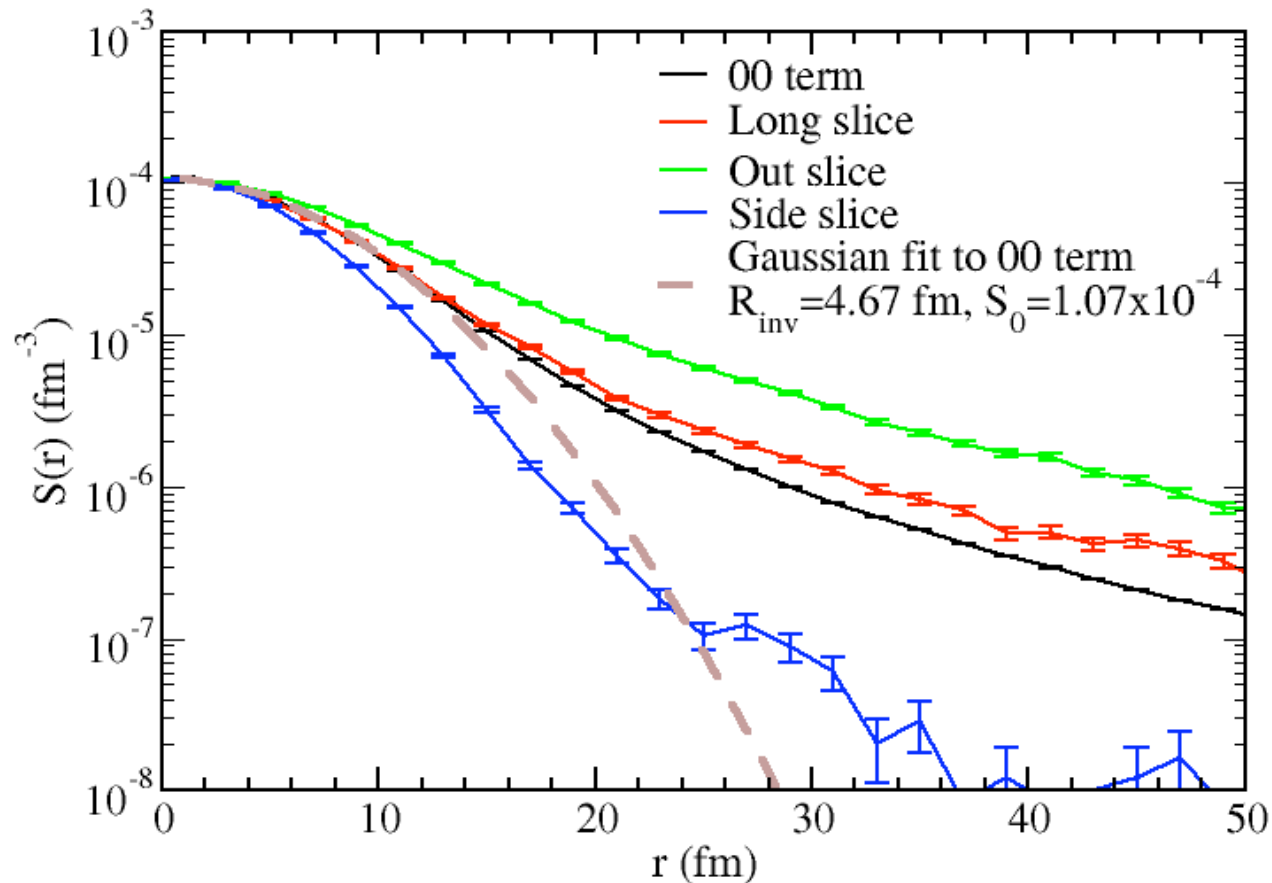
use full 3-body decay kinematics



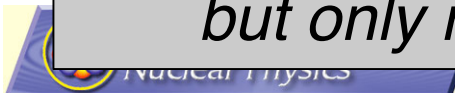
Simple Model, Interesting Results



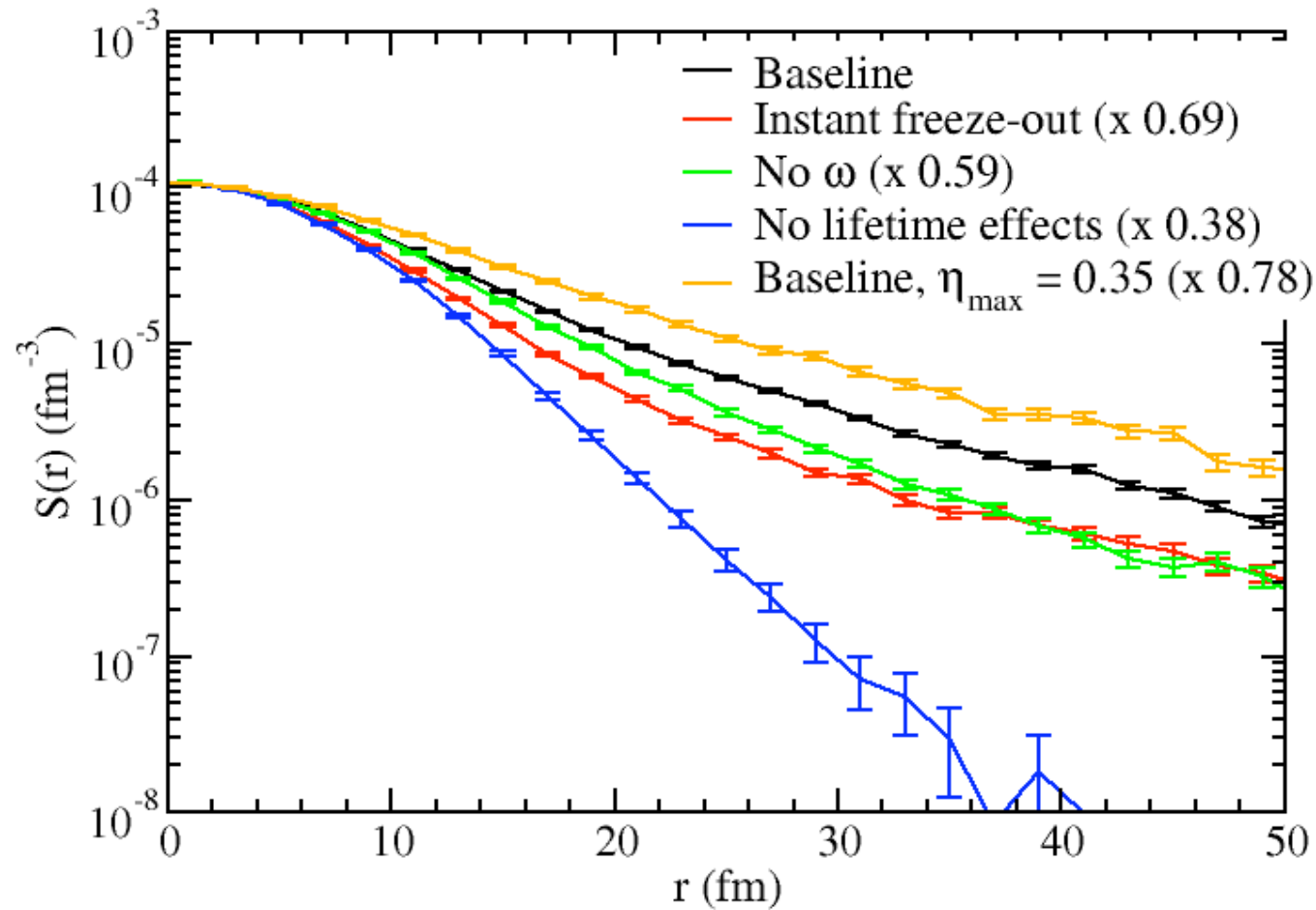
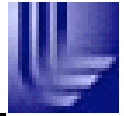
Set $R_x=R_y=R_z=4$ fm, $\tau_{f/o}=10$ fm/c, $T=175$ MeV, $f=0.56$



High velocity π 's get bigger boost. Boost + finite $\tau \Rightarrow$ tail,
but only modest core increase, in L,O directions.

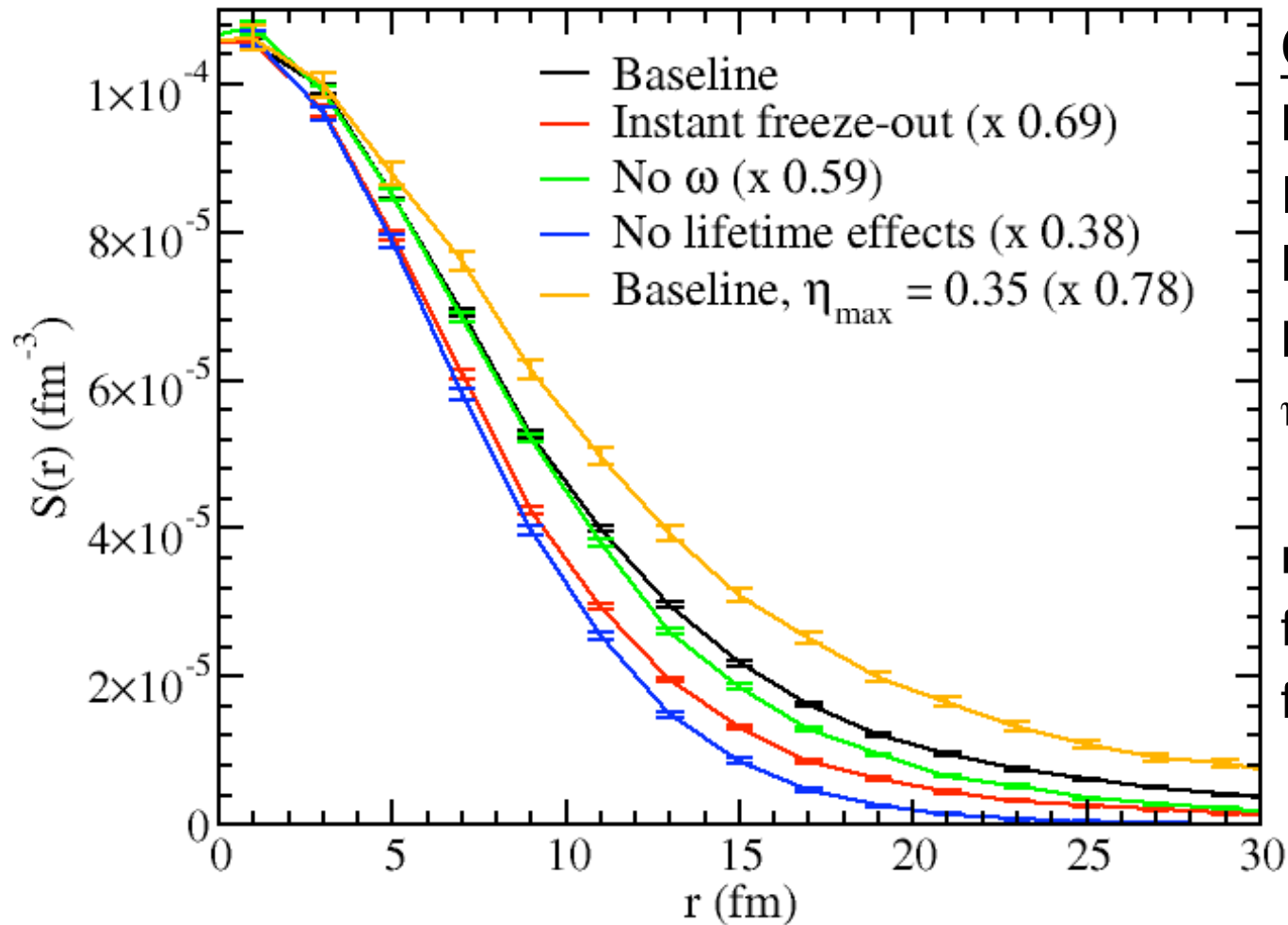
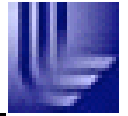


Focus on Outwards Direction



Lifetime effects have dramatic affect on non-Gaussian halo in outwards direction.

Focus on Outwards Direction



Gaussian R_0 fits:

Baseline: 5.63

Instant f/o: 4.78

No ω : 5.35

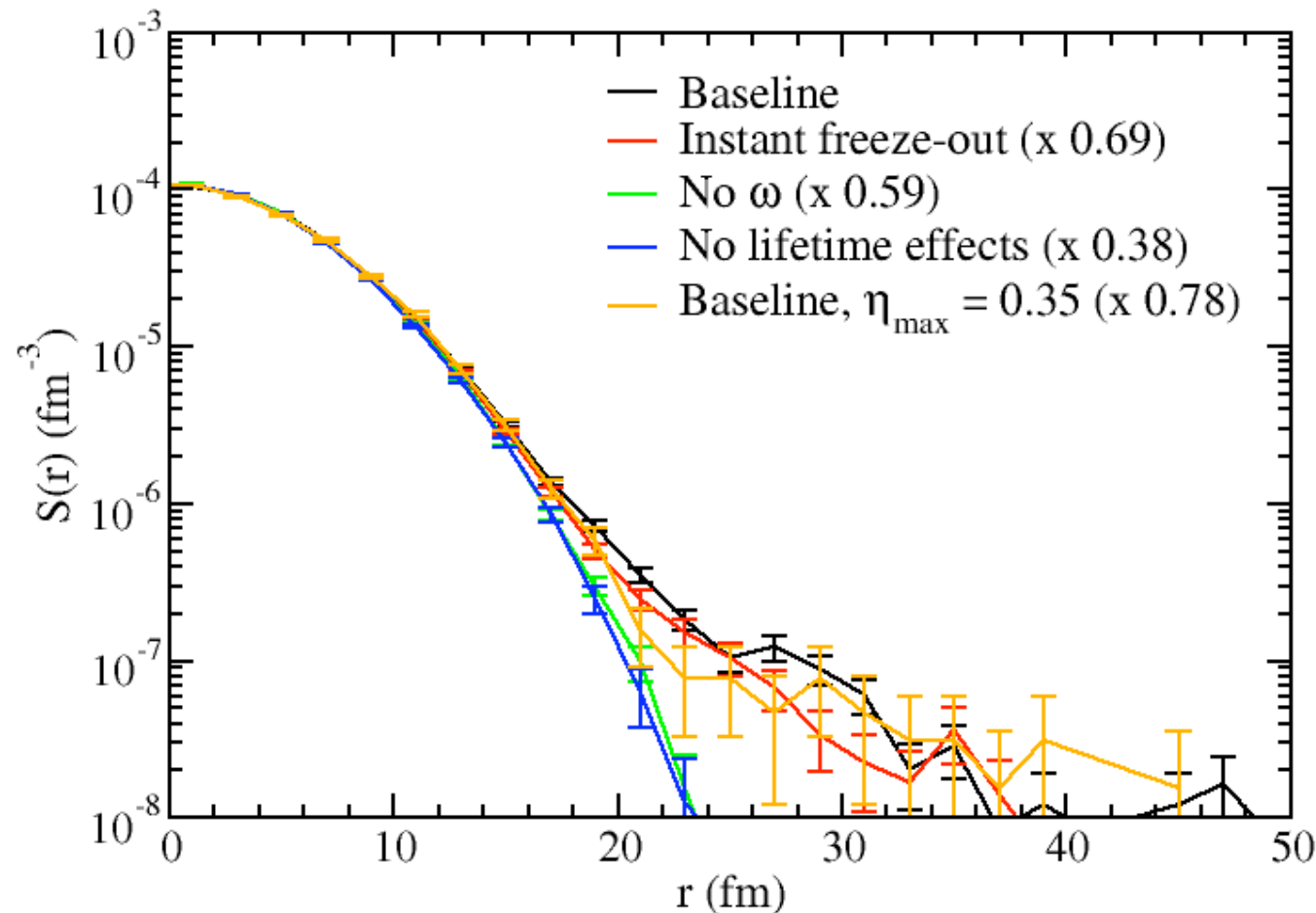
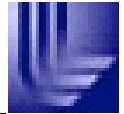
No lifetime: 4.48

$\eta_{\text{max}}=0.35$: 6.51

radii in fm,
fit w/ $r < 15$ fm,
fit tolerance 1%

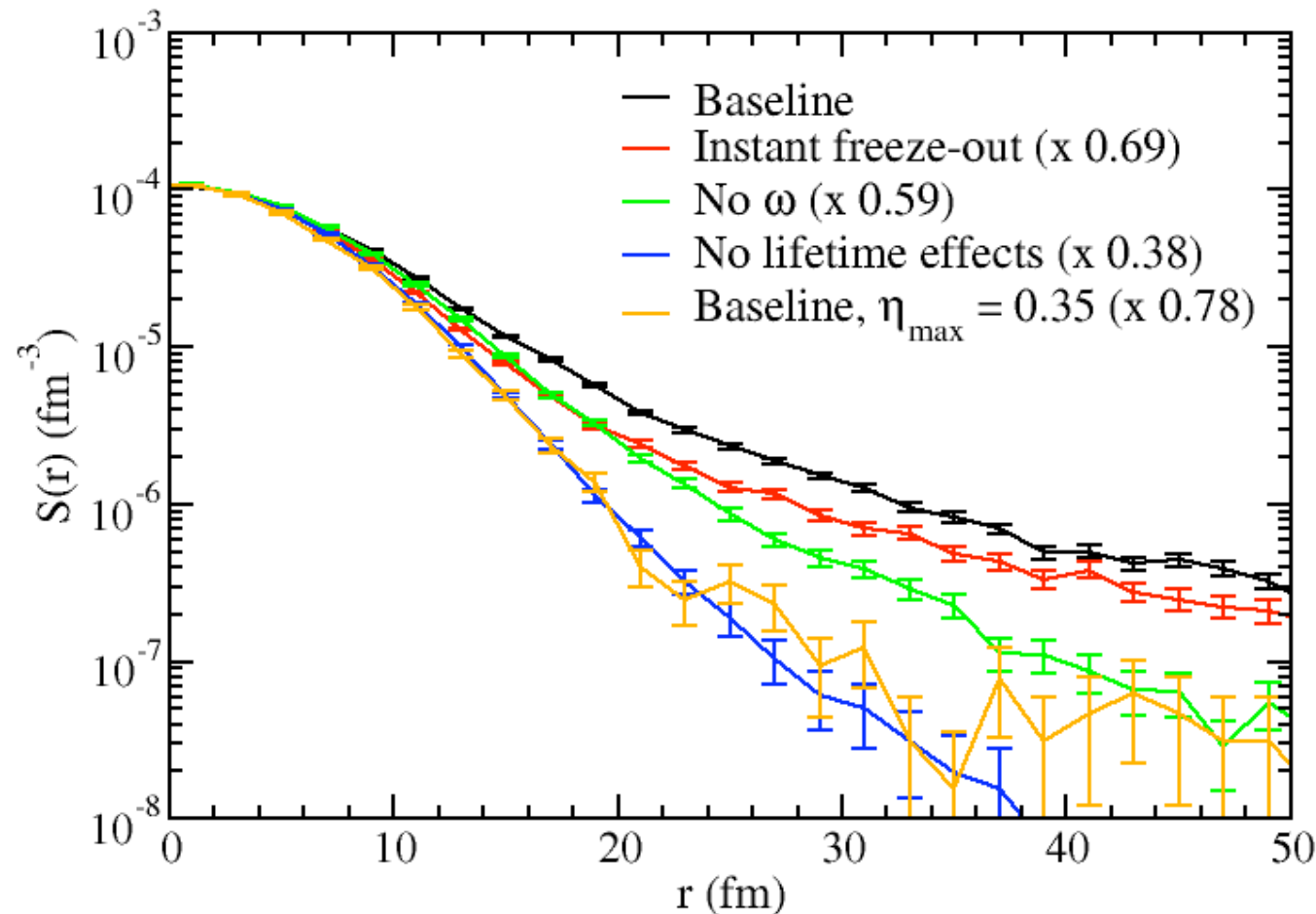
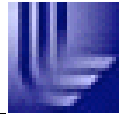
... but only *modest* core changes.

Focus on Sidewards Direction



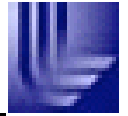
Minor changes deep in tail reflecting geometry of ω induced halo.
All curves have core radii ~ 4 fm.

Focus on Longitudinal Direction



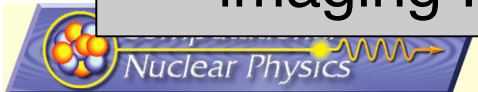
Rapidity cut eliminates tail,
otherwise similar to outwards direction.

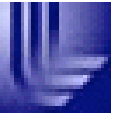
Take Home Messages from Model



- * Need emission duration + particle motion to get lifetime effect
- * Finite lifetime can create tail, w/o changing core radii substantially
- * RHIC HBT puzzle #1 could be Gaussian fit missing tail from long lifetime
- * Resonance effects detectable, but similar to freeze-out duration effects
- * Maybe need to image 3d kaon correlations to observe system lifetime & determine whether f/o is sudden or not
- * Don't be fooled by acceptance effects

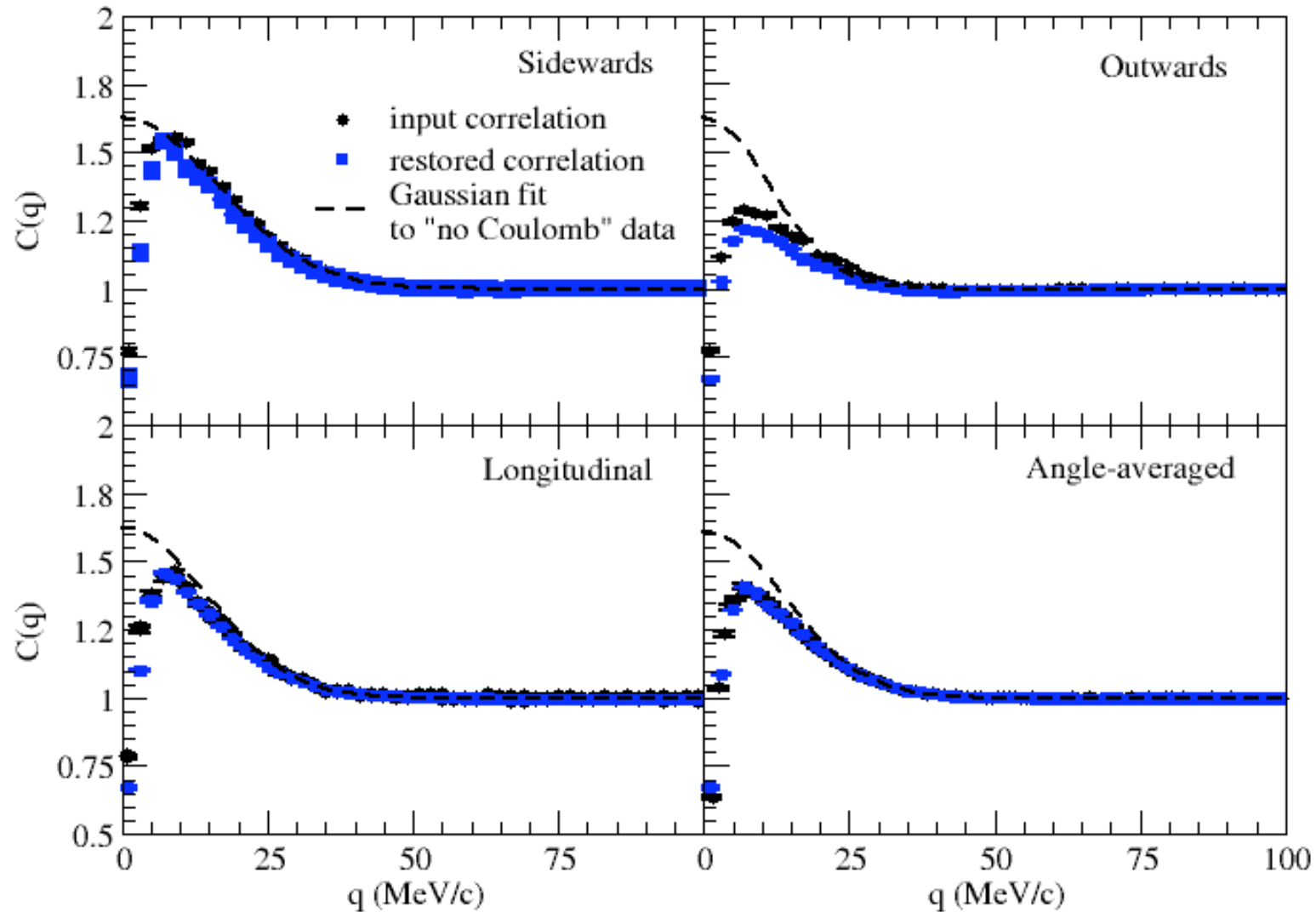
Since source tails hide in Coulomb hole of correlation, imaging RHIC data should shed light on puzzle





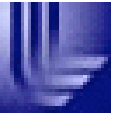
Warning:
You are about to see preliminary plots!

Correlations from Baseline model



Used S. Pratt's CRAB code

Breaking Problem into 1d Problems



Expand in Y_{lm} 's and Legendre polynomials:

$$C_{\ell m}(q) - \delta_{\ell 0} = 4\pi \int_0^\infty dr r^2 K_\ell(q, r) S_{\ell m}(r)$$

Where
$$K(\vec{q}, \vec{r}) = \sum_{\ell=0}^{\ell_{max}} (2\ell + 1) K_\ell(q, r) P_\ell(\hat{q} \cdot \hat{r})$$

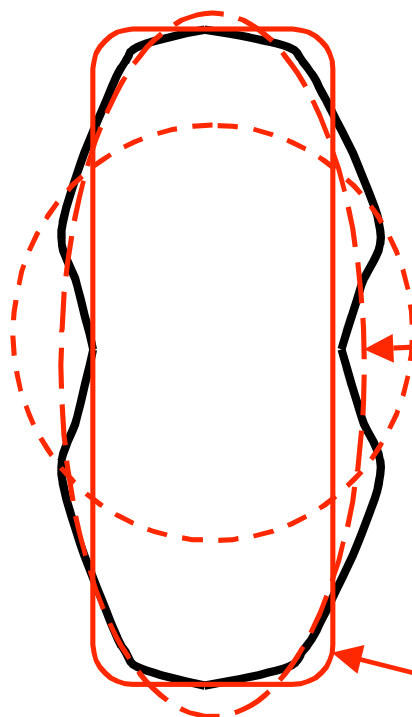
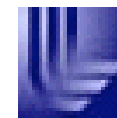
$$C(\vec{q}) = \sqrt{4\pi} \sum_{\ell=0}^{\ell_{max}} \sum_{m=-\ell}^{\ell} C_{\ell m}(q) Y_{\ell m}(\hat{q})$$

$$S(\vec{r}) = \sqrt{4\pi} \sum_{\ell=0}^{\ell_{max}} \sum_{m=-\ell}^{\ell} S_{\ell m}(r) Y_{\ell m}(\hat{r})$$



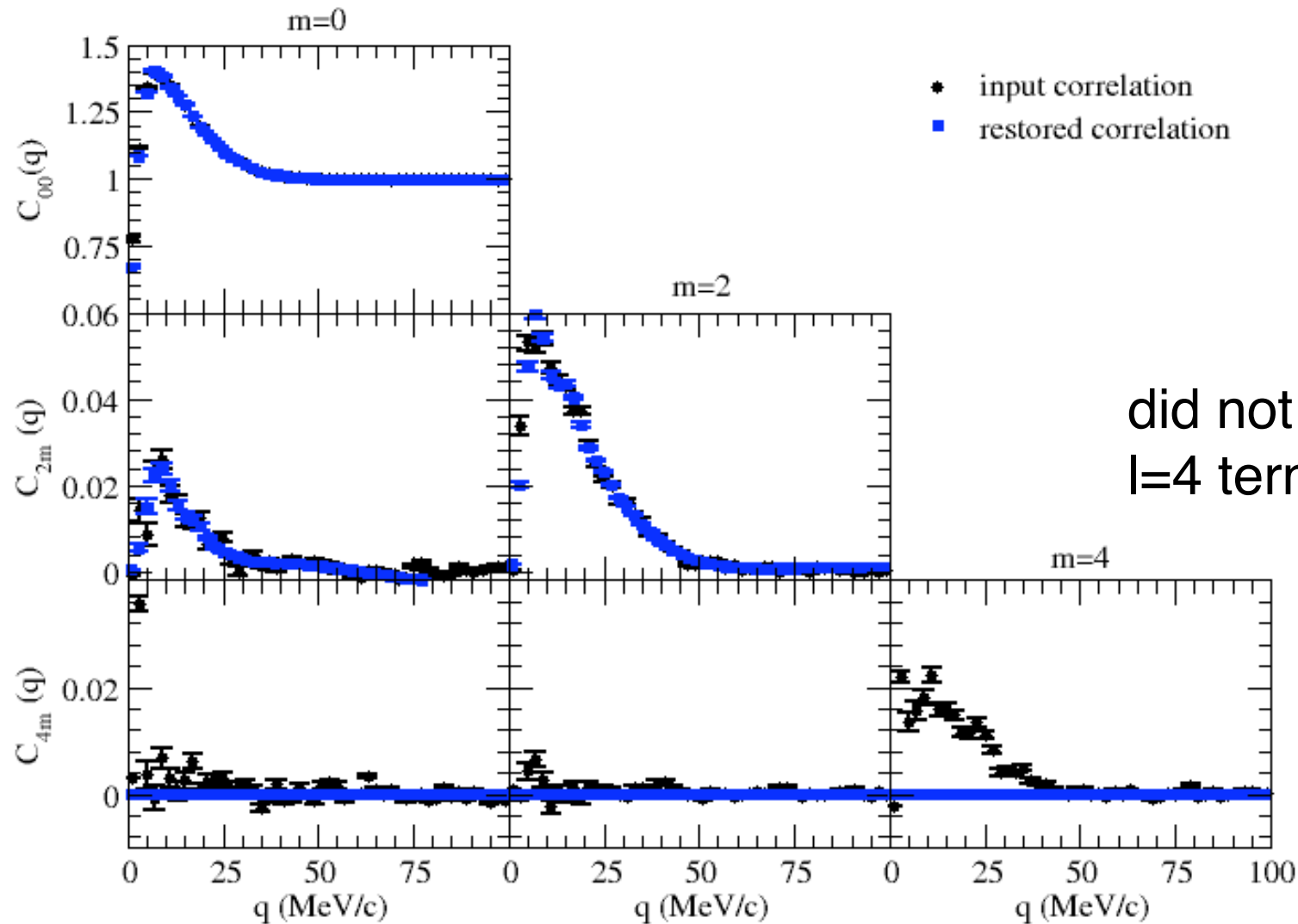
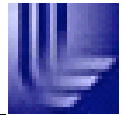
Cartesian harmonics give analogous expressions

What do the terms mean?



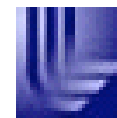
- $\ell = 0$: Angle averaged correlation, get access to R_{inv}
- $\ell = 1$: Access to Lednicky offset, i.e. who emitted first (unlike only)
- $\ell = 2$: Shape information:
access to R_O, R_S, R_L :
 $C_{20} \rightarrow R_L$
 $\alpha C_{00} - (C_{20} \pm C_{22}) \rightarrow R_S, R_O$
- $\ell = 3$: Boomerang/triaxial deformation (unlike only)
- $\ell = 4$: Squares off shape

Baseline model's terms



did not image
 $l=4$ terms yet

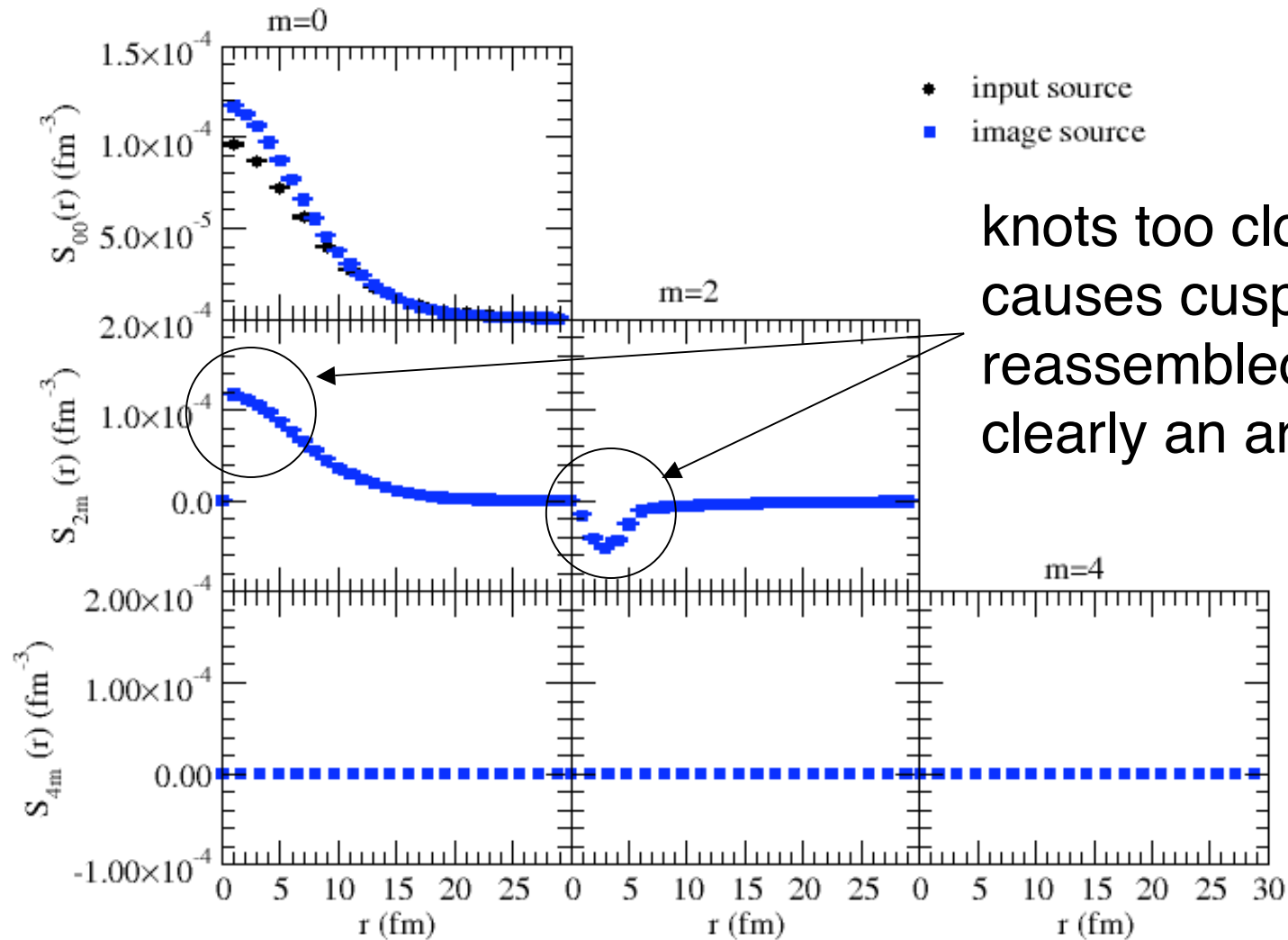
Imaging Summary



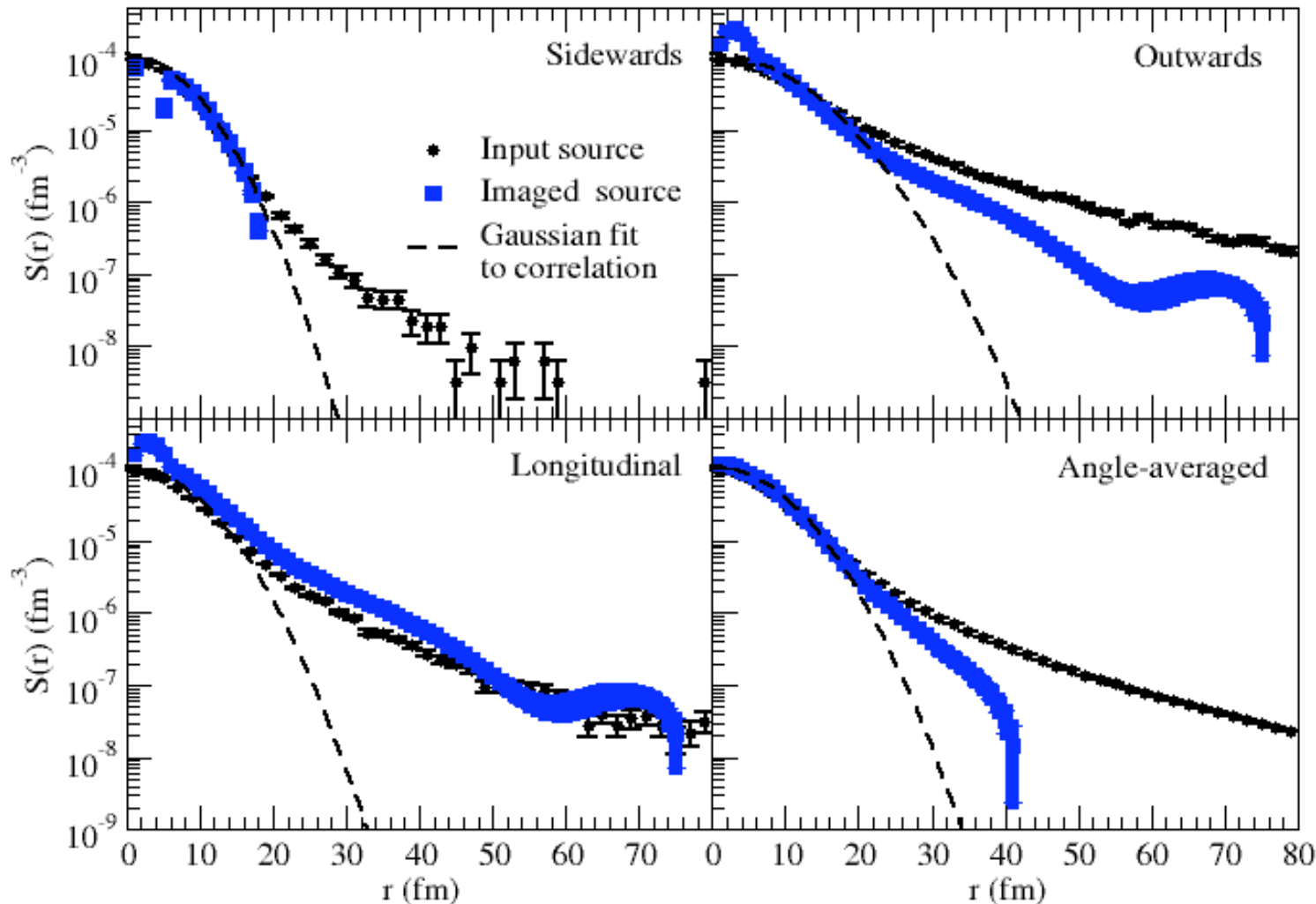
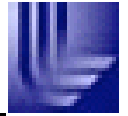
- * Use CorAL version 0.3
- * Source radial terms written in Basis Spline representation
 - knots first set using Sampling Theorem from Fourier Theory, then we optimize them to get best χ^2
 - use 3rd degree splines
- * Use full Coulomb wavefunction, symmetrized
- * Use generalized least-square for inversion
- * Use constraints to stabilize inversion
- * Cut off at finite l , q in input correlation



Terms in imaged source

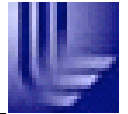


Reassembled imaged source



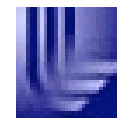
- Cusp really bad
- Tail hard to resolve
- Gaussian fit actually not bad for core
- $l_m=00$ term needs a lot of tuning ($\chi^2/\text{ndf}=8.4$)

Imaging tails in 3d is *hard*...



- * Not shown: images of other test cases -- unfortunately they don't look much different from baseline model, yet...
- * This is just first pass at imaging these test models, there is a lot of tuning to do:
 - finer q bins in correlation help a lot
 - need those $l=4$ terms
 - better statistics in model might help
 - newly coded “optimal knot” algorithm needs work
- * Can we get the tail's “time constant” from the images? Not easily, $R_{\text{halo}} \sim v\tau$, but lot's of different v 's comprise same source.

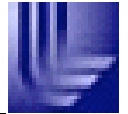
CorAL Features



- * Variety of kernels:
 - Coulomb, NN interactions, asymptotic forms
 - Any combination of p , n , π^{+-0} , K^{+-} , Λ , plus some exotic pairs
- * Fit a 1d or 3d correlation w/ variety of Gaussian sources
- * Directly image a correlation, in 1d or 3d
- * Build model correlations/sources from
 - OSCAR formatted output,
 - Blast Wave,
 - variety of simple models
- * Build model correlations/sources in Spherical or Cartesian harmonics

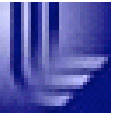


CorAL Status



- * Merging of three related projects:
 - original CorAL by M. Heffner, fitting correlations w/ source convoluted w/ full kernel
 - HBTprogs in 3d by D. Brown, P. Danielewicz, imaging sources from correlation data w/ full kernels
 - CRAB (++) by S. Pratt, use models to generate correlations, sources w/ full kernels
- * Rewritten in C++, open source, nearly stand alone (depends on GSL only)
- * Developed on MacOS X, linux, cygwin
- * Time scale for 1.0 release: end of summerish

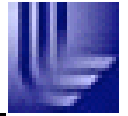




Extra Slides



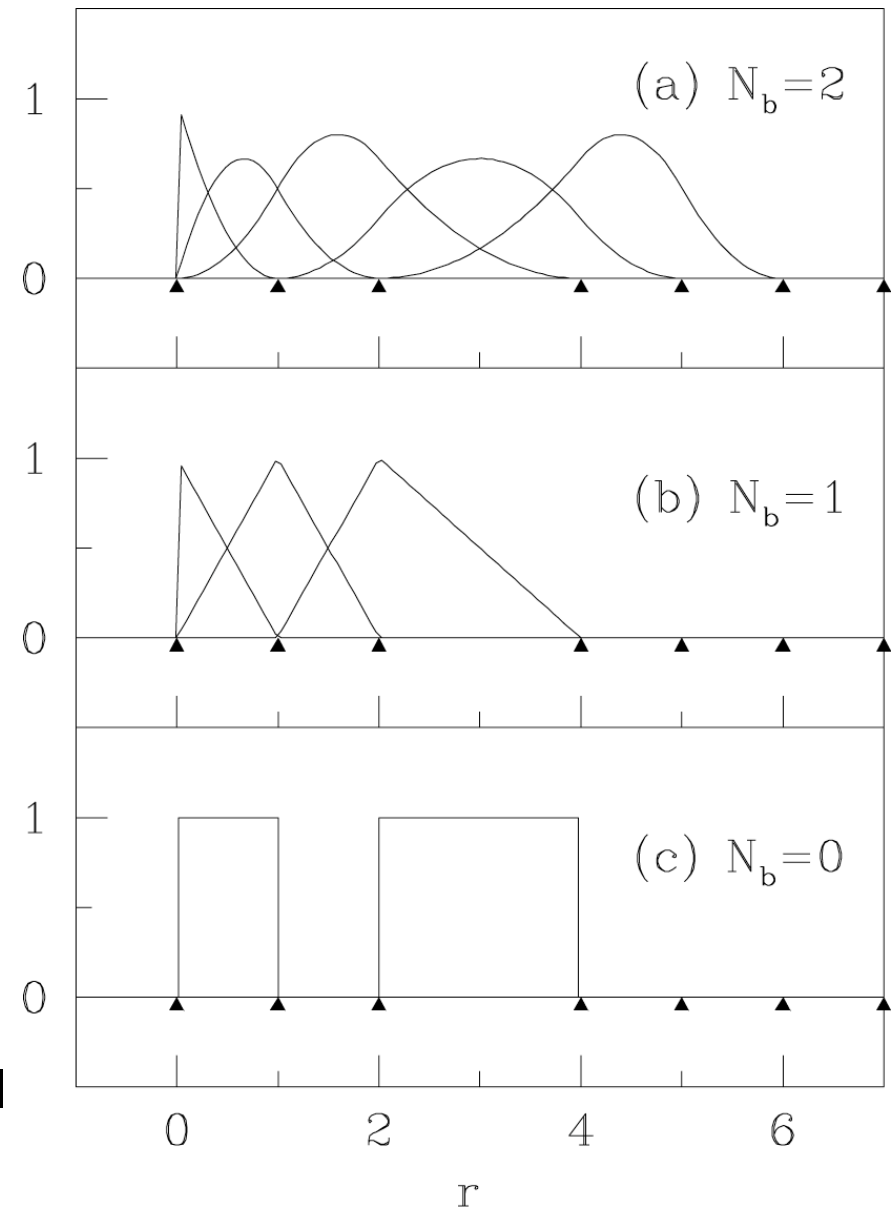
Representing the Source Function



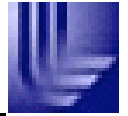
Radial dependence of each term
in terms of Basis Splines:

$$S_{\ell m}(r) = \sum_{j=1}^{N_c} S_{j\ell m} B_j(r)$$

- Generalized Splines:
 - $N_b=0$ is histogram,
 - $N_b=1$ is linear interpolation,
 - $N_b=3$ is equivalent to cubic interpolation.
- Recursion relations == fast to evaluate, differentiate, integrate
- Resolution controlled by knot placement
- Form basis for \mathbf{R}^2 , but not orthonormal



The inversion process



Koonin-Pratt eq. in matrix form: $\mathbf{C} = \mathbf{K} \cdot \mathbf{S}$

this is an ill-posed problem

- kernel not square
- kernel possibly singular
- noisy data
- error propagation

Practical solution to linear inverse problem, minimize χ^2 :

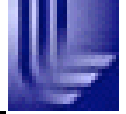
$$\chi^2 = \sum_i \frac{(C_i - \sum_j K_j S_j)^2}{\Delta^2 C_{ii}}$$

Most probable source is: $\mathbf{S} = \Delta^2 \mathbf{S} \cdot \mathbf{K}^T \cdot (\Delta^2 \mathbf{C})^{-1} \cdot \mathbf{C}^{\text{obs}}$

With covariance matrix: $\Delta^2 \mathbf{S} = (\mathbf{K}^T \cdot (\Delta^2 \mathbf{C})^{-1} \cdot \mathbf{K})^{-1}$



Constraints for 3d problem(s)



Inversion can be stabilized with constraints.

Constraints we use:

Constraint	Purpose	Functional form
$r = 0$ is a maximum of $S(\mathbf{r})$ (for like pairs only)	Constrain the higher ℓ components that are not well controlled due to the r^ℓ dependence of terms in the spherical harmonic expansion.	$\frac{\partial S_{\ell m}}{\partial r}(r \rightarrow 0) = 0 \quad \forall \ell, m$ $S_{\ell m}(r \rightarrow 0) = 0 \quad \forall \ell, m, \ell \neq 0$
$S(\mathbf{r}) = 0$ at $r = r_{\max}$	Smooth oscillations in the source at high \mathbf{r} caused by aliasing of statistical and experimental noise in correlation.	$S_{\ell m}(r_{\max}) = 0 \quad \forall \ell, m$
$S(\mathbf{r}) = 0$ is flat as $r \rightarrow r_{\max}$	Smooth oscillations in the source at high \mathbf{r} caused by aliasing of statistical and experimental noise in correlation.	$\frac{\partial S_{\ell m}}{\partial r}(r_{\max}) = 0 \quad \forall \ell, m$

TABLE I: Equality constraints on the Basis Spline representation for non-spherically symmetric sources.