WPCF, Kroměříž, August 2005

# **Non-Gaussian Effects in Identical Pion Correlation Function at STAR**

Michal Bysterský

August 16, 2005

### Outline

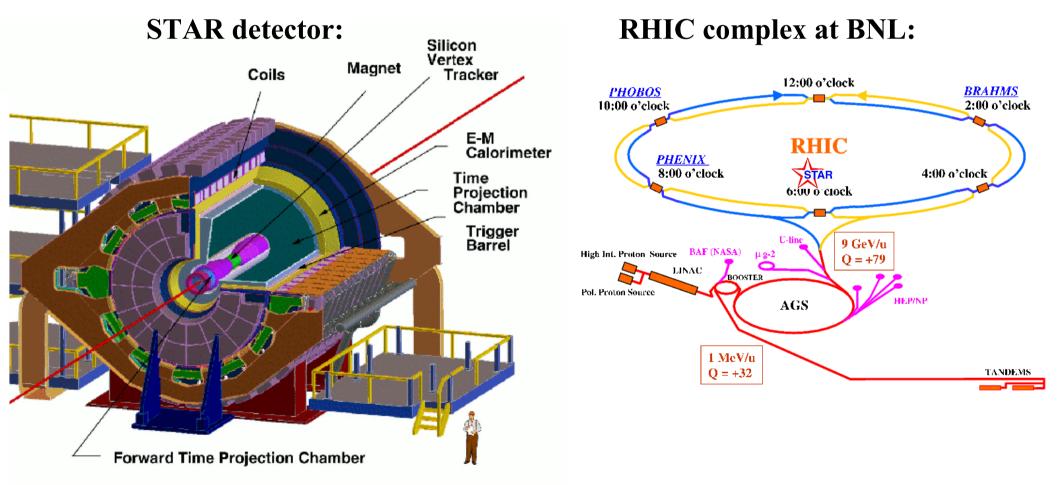
- Motivation
- STAR data
- Event and Track selection
- Identical  $\pi$ - $\pi$  correlation function
- Bowler-Sinyukov fit to data
- Levy source distribution fit
- Edgeworth expansion
- Summary

- Why do we care about non-Gaussian issue?
  - Source distribution function is in most models non-Gaussian and standard methods of fitting experimental CF assume Gaussian source.
  - Need to parametrize source properly in order to minimize systematic errors.
- Possible methods of studying the non-Gaussian effects of CF include:
  - Source imaging, see talk by P.Danielewicz, P.Chung, D.Brown
    - Spherical harmonics, see talk by Z.Chajęcki
  - Levy stable source distribution, see talk by T.Csörgő
  - Edgeworth expansion

## STAR data

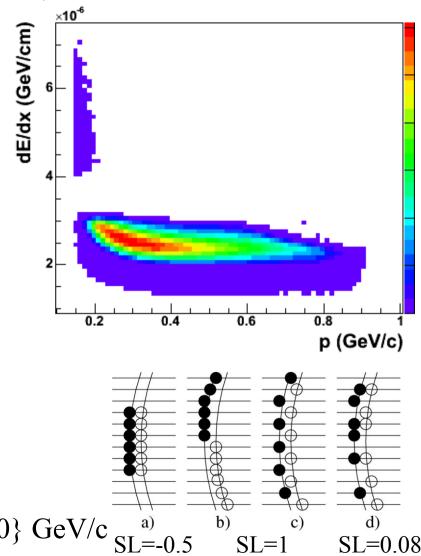
- Au+Au collisions at energy  $\sqrt{s_{_{NN}}} = 200 \text{ GeV}$
- Year 2004 data, Full Field (0.5 T)
- ~11 M MinBias events



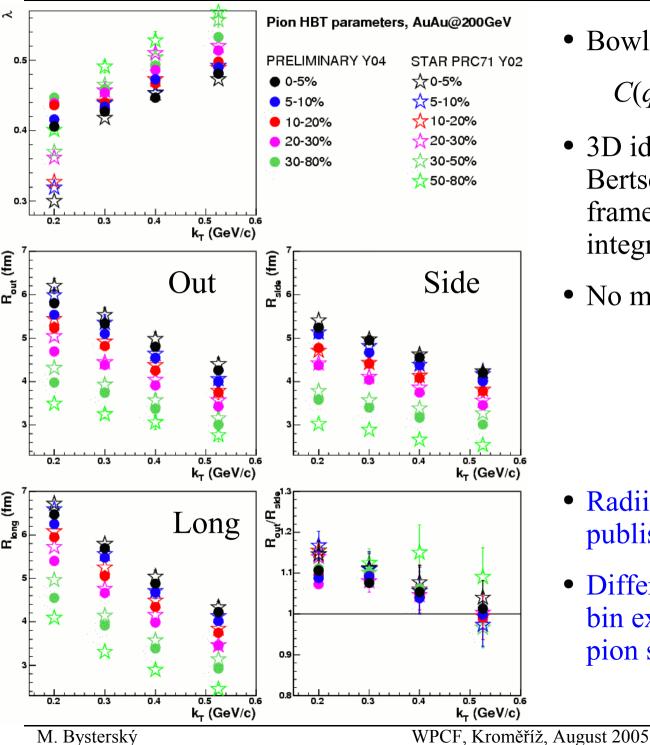


#### **Event and Track selection**

- Same cuts as in STAR, Phys. Rev. C 71 (2005) 044906
- Event cuts:
  - Centrality binning {0-5, 5-10, 10-20, 20-30, 30-80} %
  - $zVertex \pm 25 cm$
- Track cuts:
  - pion dE/dx band  $\pm 2$  s
  - remove dE/dx electron band
  - $p_{\rm T} = \{0.15, 0.80\}$  GeV/c
  - $y = \{-0.5, 0.5\}$
- Pair cuts:
  - Id:  $\pi^+ \pi^+$ ,  $\pi^- \pi^-$
  - anti-splitting (-0.5 < SL < 0.6)
  - anti-merging (max. 5 % merged)
  - $k_{\rm T} = \{0.15 0.25, 0.25 0.35, 0.35 0.45, 0.45 0.60\}$  GeV/c <sub>SI</sub>



#### **Comparison to published STAR data**



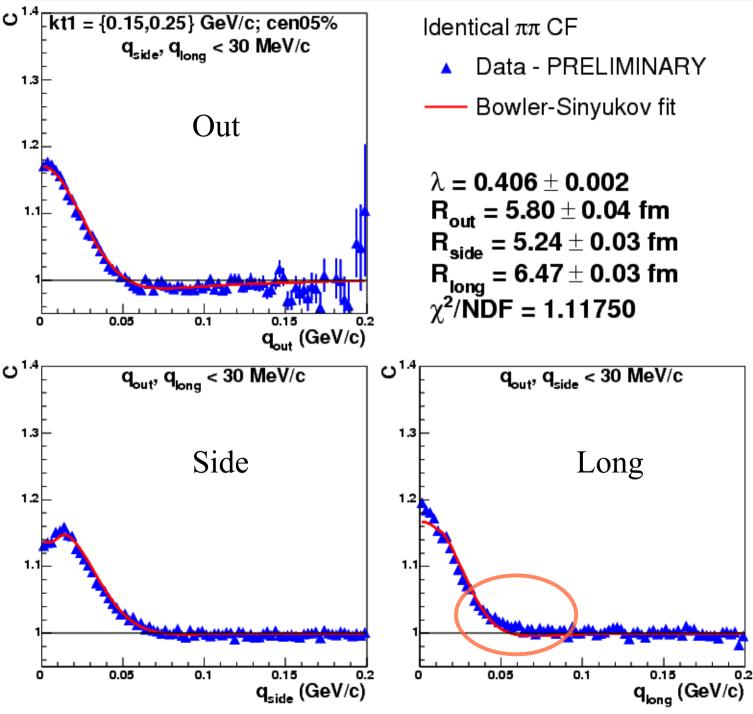
• Bowler-Sinyukov fit to data

 $C(q) = (1-\lambda) + \lambda K_{c}(1 + \exp(-\sum R_{ij}^{2}q_{i}q_{j}))$ 

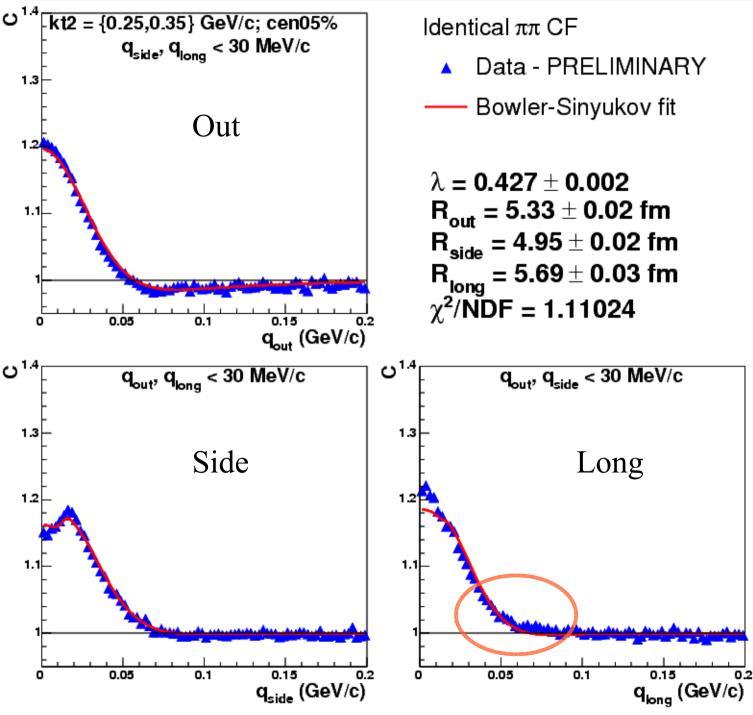
- 3D identical pion CF is fit using the Bertsch-Pratt parametrization in LCMS frame without crossterms in asimuthally integrated analyses
- No momentum resolution correction yet

- Radii are consistent within errors with published STAR PRC71 data
- Difference in lambda in the lowest kT bin explained by improved purity of pion sample

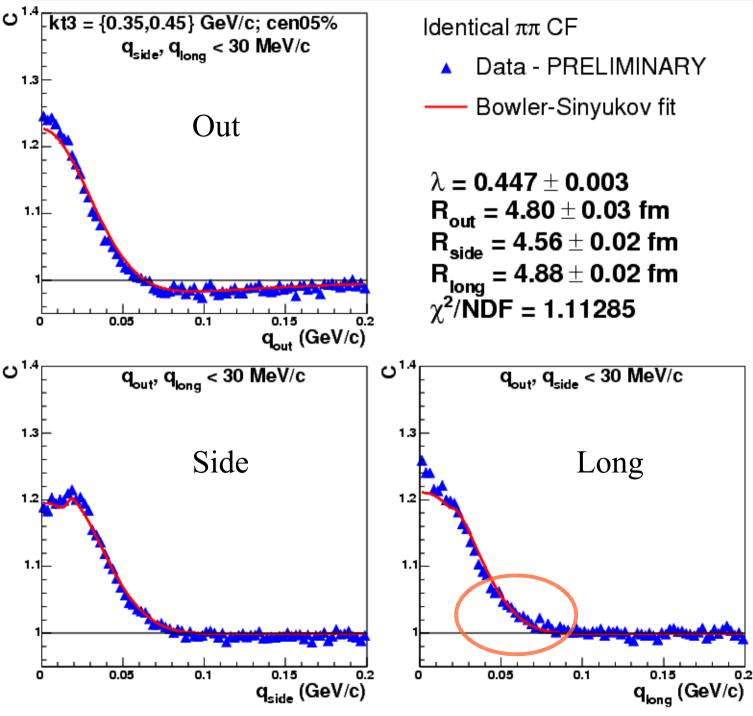
6



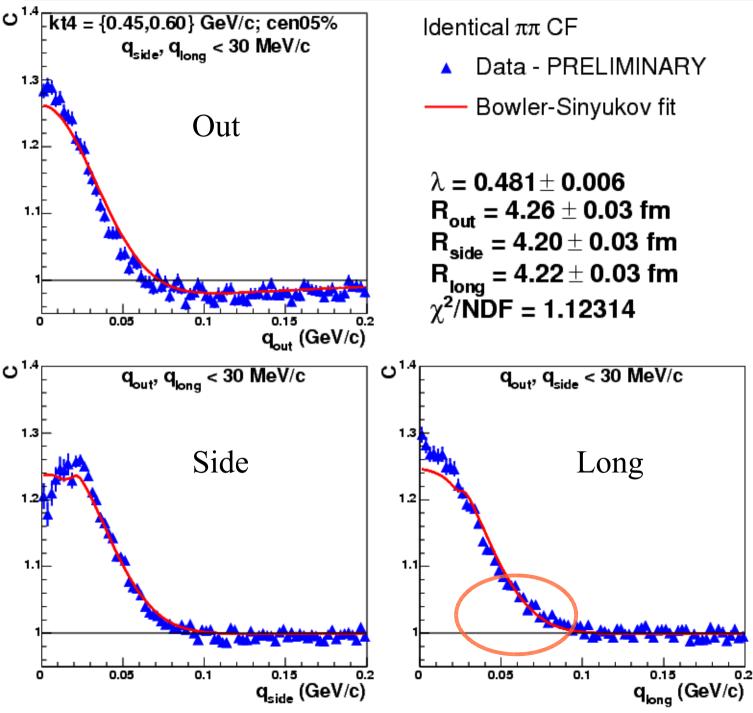
- All three radii  $R_{0,s,l}$  decrease with  $k_{T}$
- Non-Gaussian shape mostly visible in long direction



- All three radii  $R_{0,s,l}$  decrease with  $k_{T}$
- Non-Gaussian shape mostly visible in long direction



- All three radii  $R_{0,s,l}$  decrease with  $k_{T}$
- Non-Gaussian shape mostly visible in long direction



•  $\lambda$  increases with  $k_{\rm T}$ 

• All three radii  $R_{osl}$ 

decrease with  $k_{\rm T}$ 

direction

• Non-Gaussian shape

mostly visible in long

### Levy source distribution fit

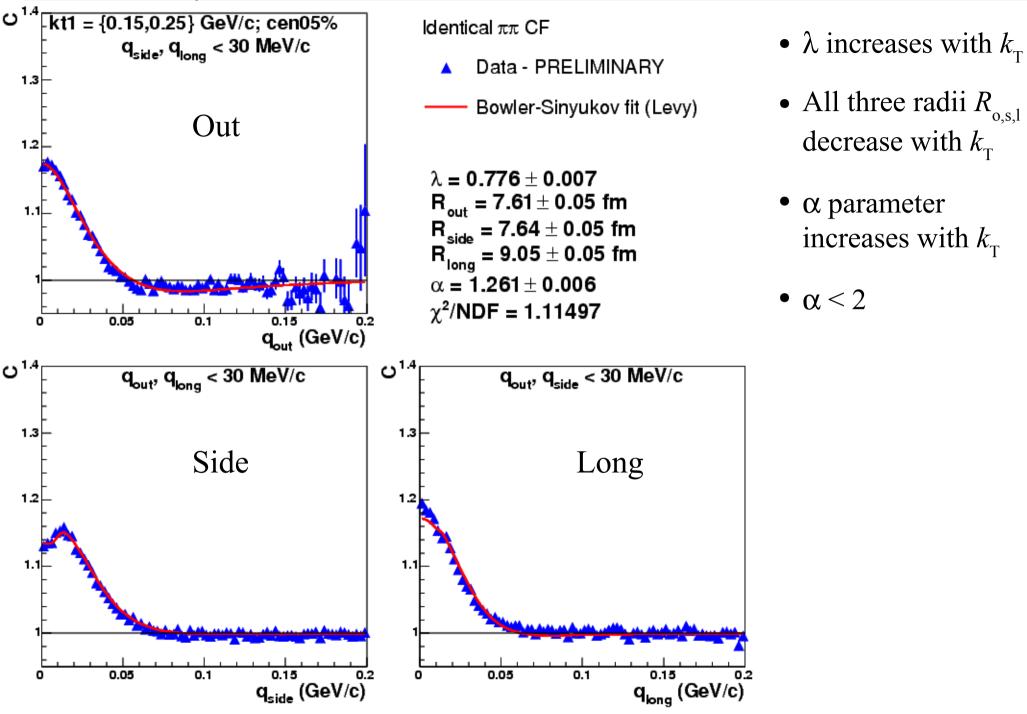
- T.Csörgő, *et al.*: Bose-Einstein correlations for Levy stable source distributions, Eur.Phys.J. C36(2004)67
- The general form of two-particle BECF

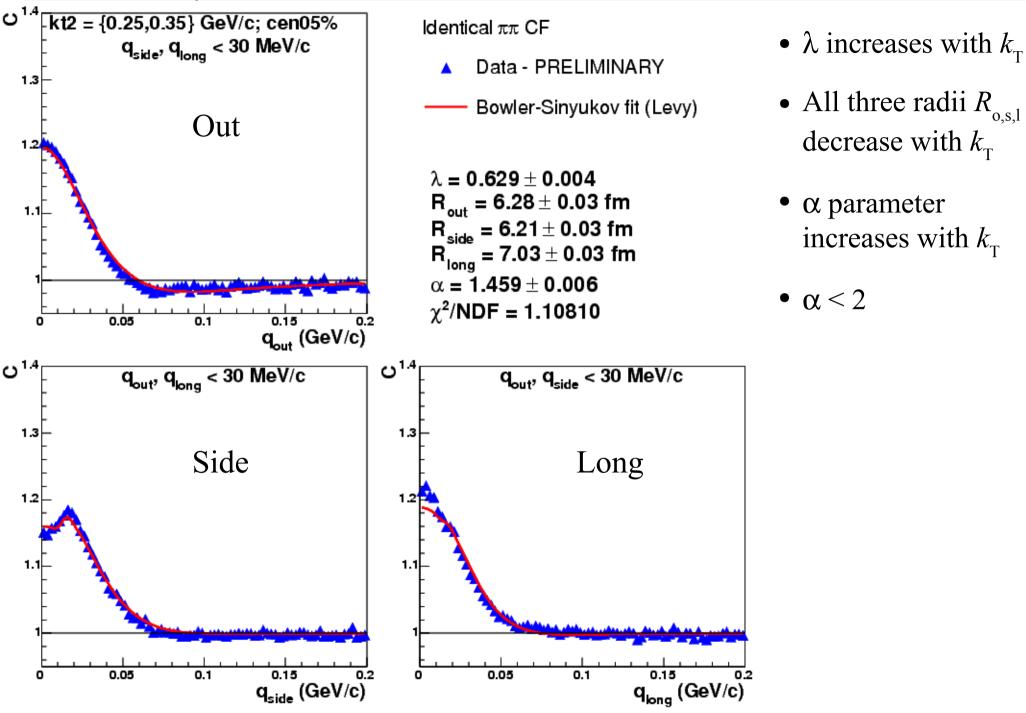
 $C(q) = 1 + \lambda \exp(-(\sum R_{ij}^2 q_i q_j)^{\alpha/2})$ 

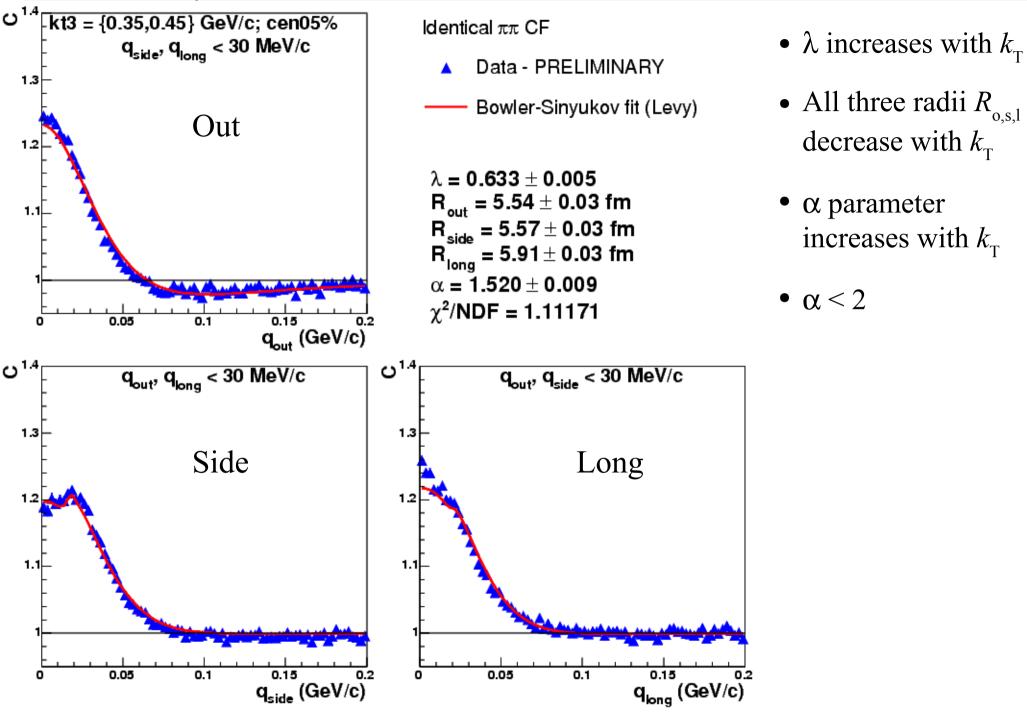
- $0 < \alpha \le 2$  ...Levy index of stability
- $\alpha < 2$  ... CF becomes more peaked than a Gaussian and it develops longer tails
- Taking into account the Coulomb effect, Levy source distribution fit to data

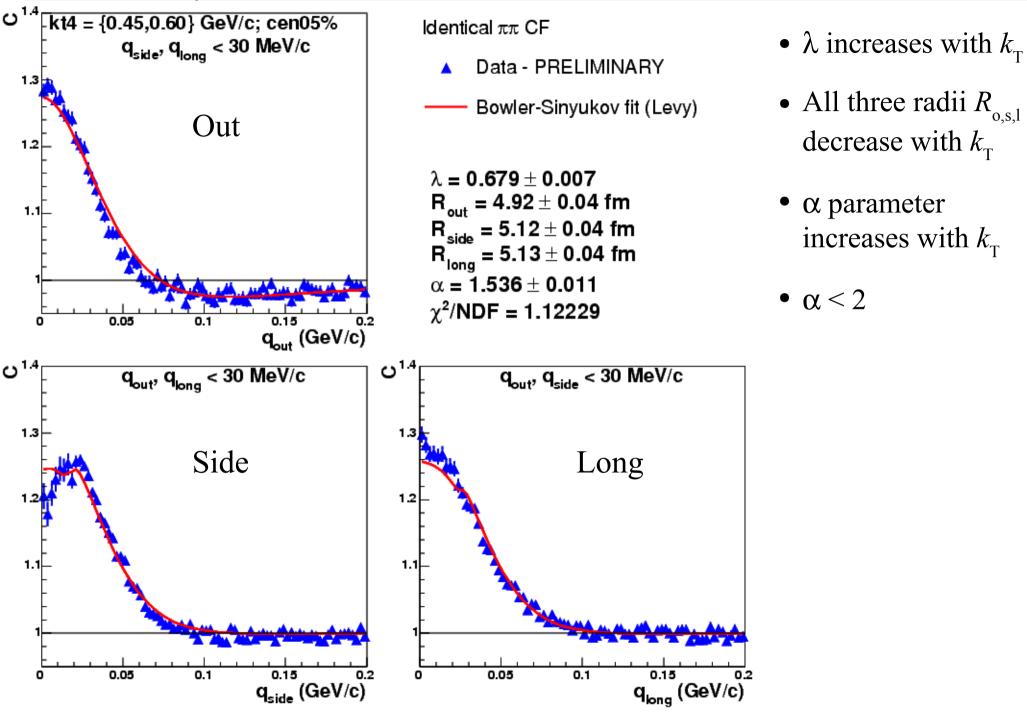
 $C(q) = (1-\lambda) + \lambda K_{c}(1 + \exp(-(\sum R_{ij}^{2}q_{i}q_{j})^{\alpha/2}))$ 

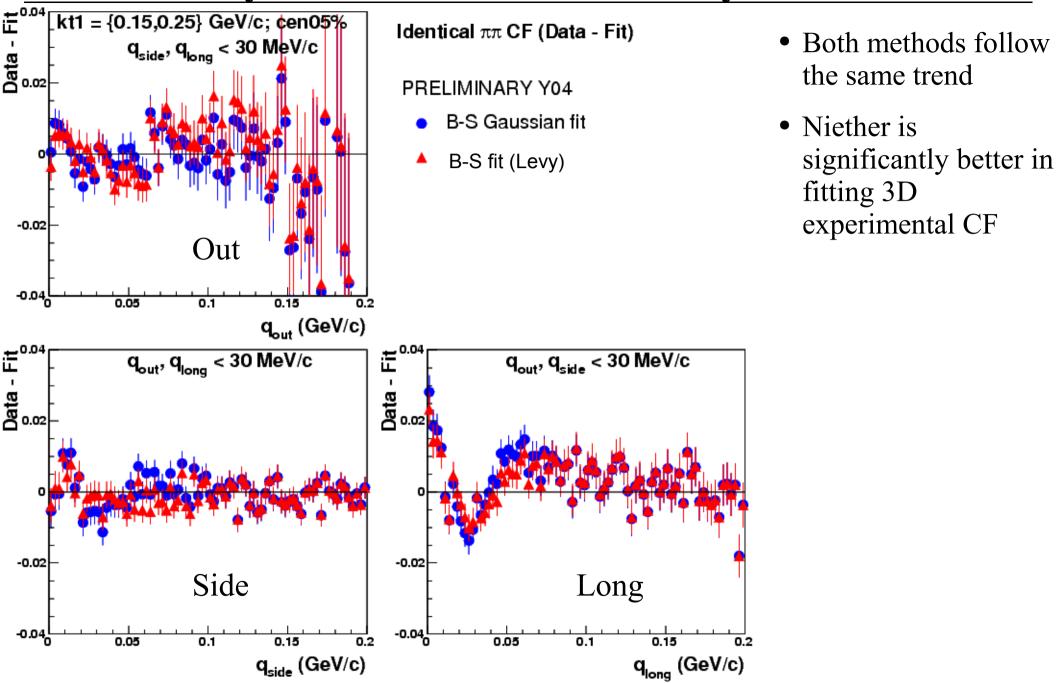
• Using Bertsch-Pratt parametrization in LCMS frame, asimuthally integrated analyses,  $R_{ii}=0$  i $\neq j$ 

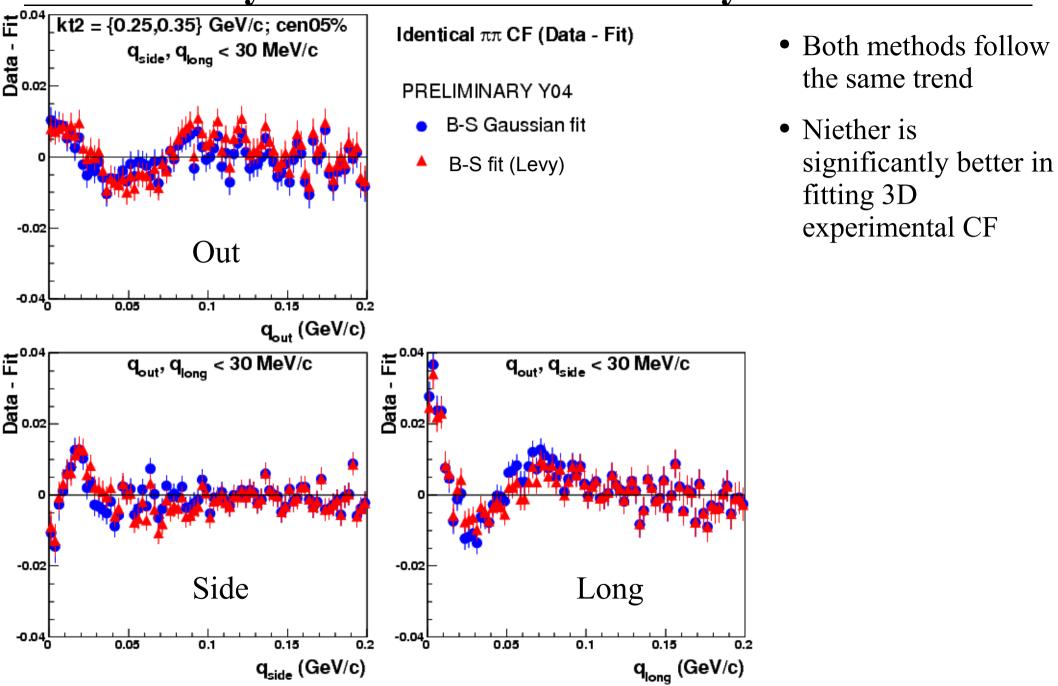


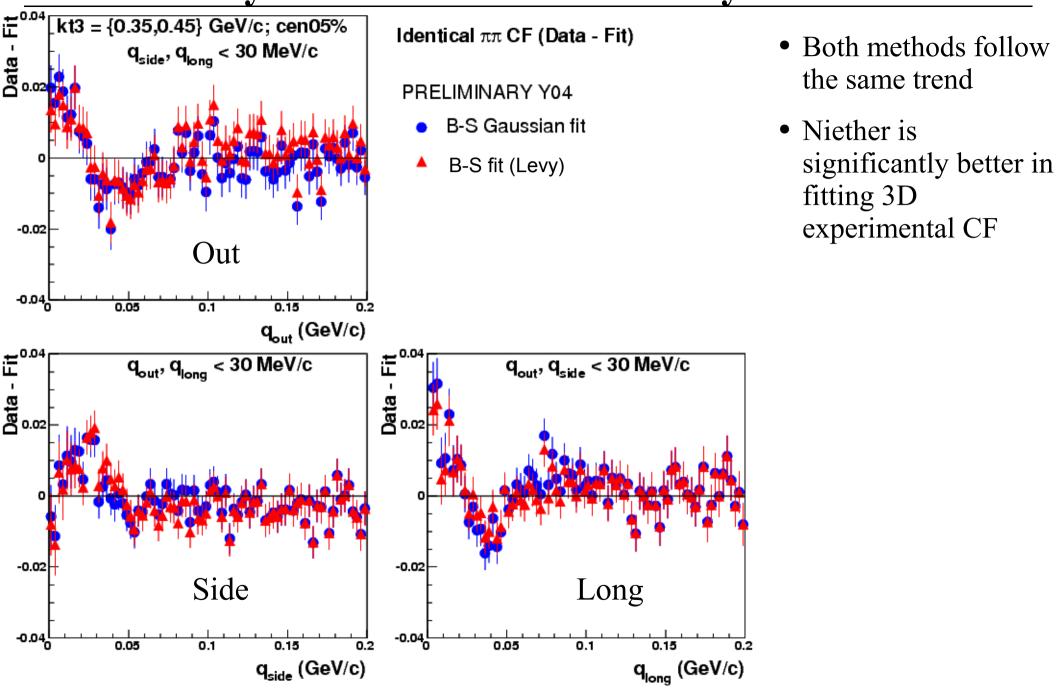


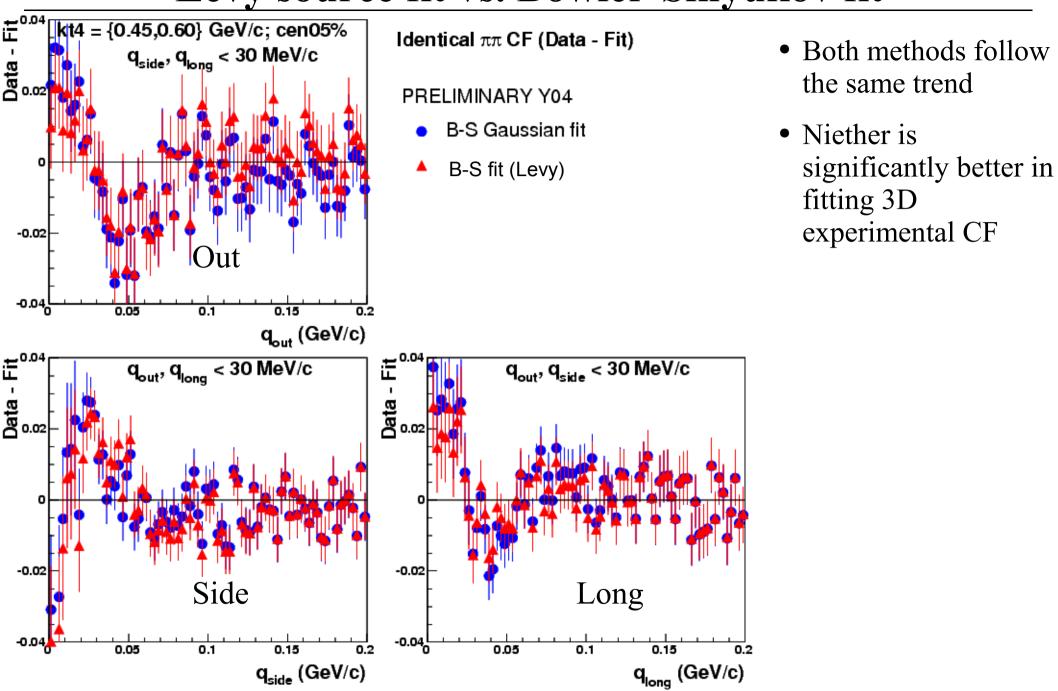






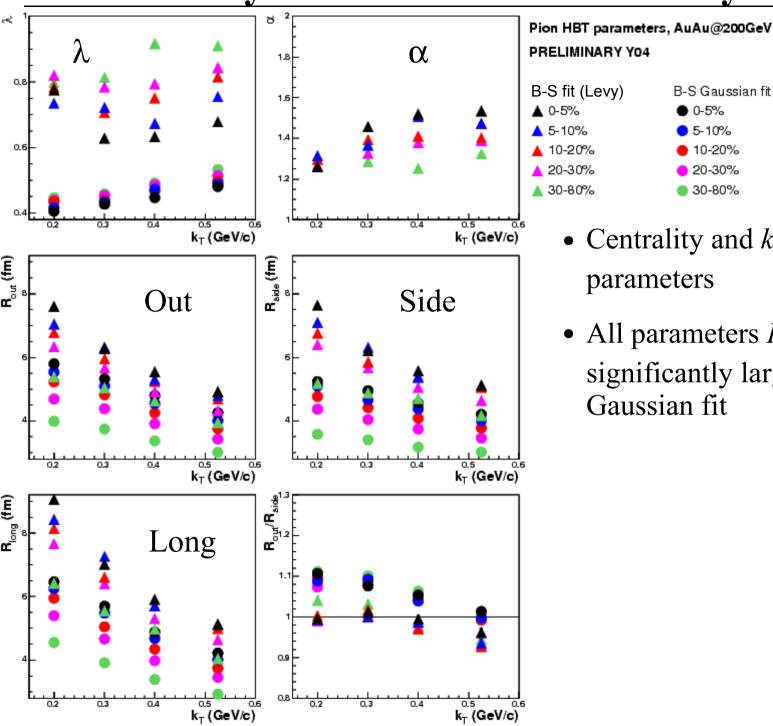






5-10%

10-20%



parameters • All parameters  $R_{0.s.1}$  and  $\lambda$  are significantly larger when compared to Gaussian fit

• Centrality and  $k_{T}$  dependence of fit

B-S Gaussian fit

0-5%

5-10%

10-20%

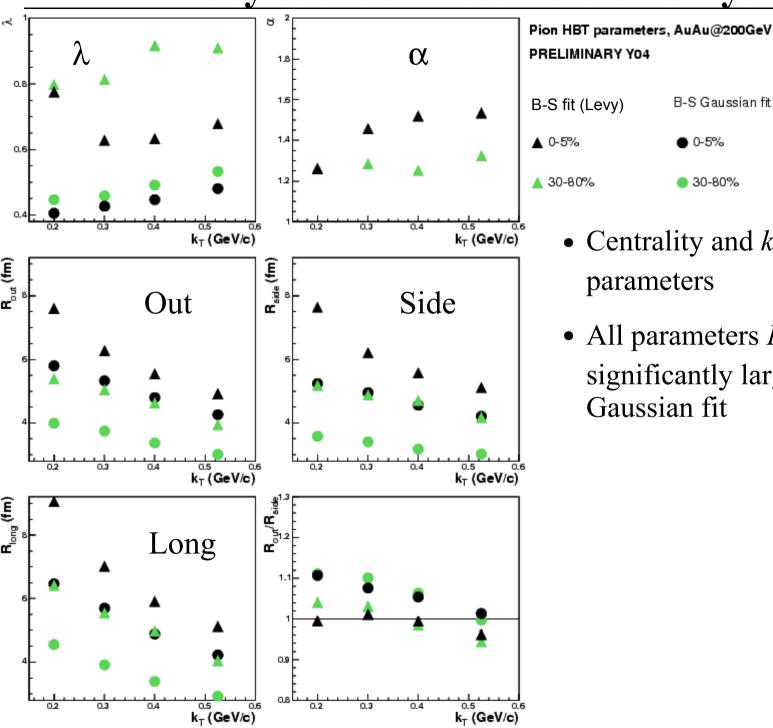
20-30%

30-80%

20

M. Bysterský

WPCF, Kroměříž, August 2005



M. Bysterský

- PRELIMINARY Y04 B-S Gaussian fit B-S fit (Levy) ▲ 0-5% 0-5% 30-80% 30-80%
  - Centrality and  $k_{T}$  dependence of fit parameters
  - All parameters  $R_{0.s.l}$  and  $\lambda$  are significantly larger when compared to Gaussian fit

- Method suggested by T.Csörgő, *et al.*, Phys.Lett. B489(2000)15, to study deviations from Gaussian CF
- Edgeworth expansion arround 3D Gaussian in B-S procedure

$$C(q_o, q_s, q_l) = (1 - \lambda) + \lambda K_{\text{coul}}(q_{\text{inv}}) + \lambda K_{\text{coul}}(q_{\text{inv}}) \cdot e^{-q_o^2 R_o^2 - q_s^2 R_s^2 - q_l^2 R_l^2} \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{o,n}}{n!(\sqrt{2})^n} H_n(q_o R_o)\right] \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{s,n}}{n!(\sqrt{2})^n} H_n(q_s R_s)\right] \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{l,n}}{n!(\sqrt{2})^n} H_n(q_l R_l)\right],$$

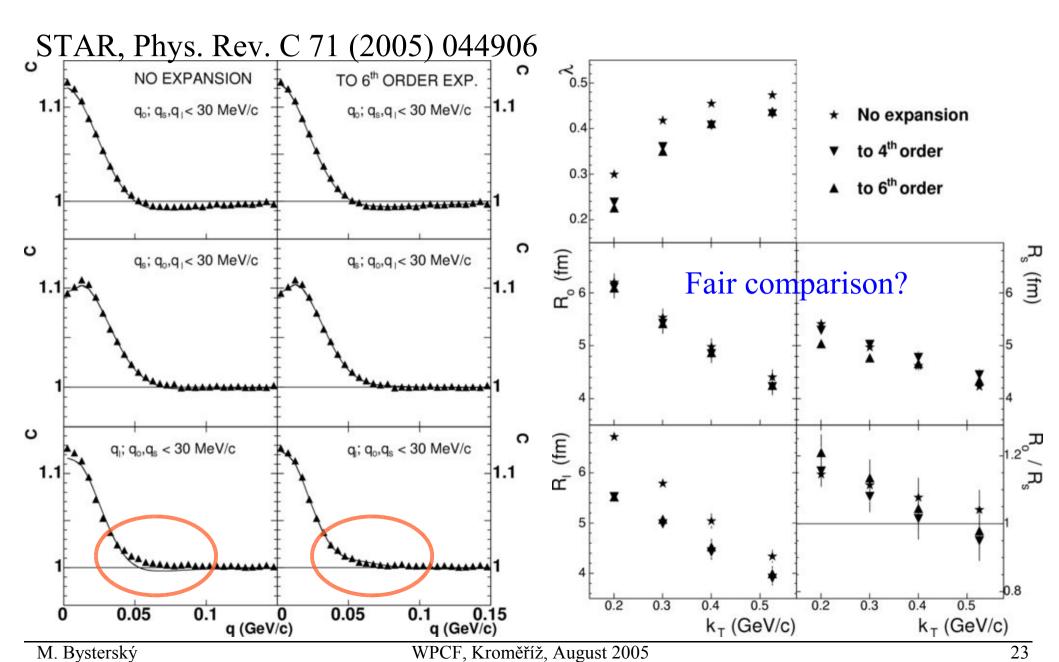
• Unable to find the physical interpretation of the fit parameters, it is not clear how to compare extracted parameters to models that assume Gaussian CF

#### Numbr of parameters

Gaussian	4
4th order	7
6th order	10

#### Edgeworth expansion fit to identical $\pi$ - $\pi$ CF

• Expansion up to 6th order is sufficient.



#### Summary

- Large statistics of year 2004 data ~11 M MinBias Au+Au events at  $\sqrt{s_{_{NN}}} = 200$  GeV, is being processed with the extraction of the  $k_{_{T}}$  dependence of interferometry parameters for 5 centrality bins.
- Gaussian fit to experimental CF is consistent within errors with STAR published data.
- Levy source distribution does not fit the experimental CF significantly better when compared to standard Gaussian fit.
- Edgeworth expansion is an improvement but there is no clear interpretation of higher order fit parameters yet.
- It seems that at RHIC Au+Au collisions Gaussian parameterization is sufficient to represent experimental CF.

#### Thank you

#### **Extra slides**

#### **Edgeworth expansion**

- Method suggested by T.Csörgő, *et al.*, Phys.Lett. B489(2000)15, to study deviations from Gaussian CF
- Edgeworth expansion arround 3D Gaussian in B-S procedure

$$C(q_o, q_s, q_l) = (1 - \lambda) + \lambda K_{\text{coul}}(q_{\text{inv}}) + \lambda K_{\text{coul}}(q_{\text{inv}}) \cdot e^{-q_o^2 R_o^2 - q_s^2 R_s^2 - q_l^2 R_l^2} \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{o,n}}{n!(\sqrt{2})^n} H_n(q_o R_o)\right] \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{s,n}}{n!(\sqrt{2})^n} H_n(q_s R_s)\right] \times \left[1 + \sum_{n=4, n \text{ even}}^{\infty} \frac{\kappa_{l,n}}{n!(\sqrt{2})^n} H_n(q_l R_l)\right],$$

• Unable to find the physical interpretation of the fit parameters, it is not clear how to compare extracted parameters to models that assume Gaussian CF

#### STAR, Phys. Rev. C 71 (2005) 044906

/	2		、 <i>,</i>	
$k_T \; (MeV/c)$	150 - 250	250 - 350	350 - 450	450-600
$\lambda$	$0.30\pm0.01$	$0.42\pm0.01$	$0.45\pm0.01$	$0.47\pm0.01$
$\lambda (4^{\text{th}} \text{ ord.})$	$0.24\pm0.01$	$0.36\pm0.01$	$0.41\pm0.01$	$0.43\pm0.01$
$\lambda~(6^{\rm th}~{\rm ord.})$	$0.23\pm0.01$	$0.35\pm0.01$	$0.41\pm0.01$	$0.44\pm0.01$
$R_o$	$6.16\pm0.01$	$5.51\pm0.01$	$4.88\pm0.02$	$4.32\pm0.02$
$R_o$ (4 <sup>th</sup> ord.)	$6.07\pm0.04$	$5.40\pm0.03$	$4.75\pm0.03$	$4.14\pm0.04$
$\kappa_{o,4}$	$0.37\pm0.05$	$0.36\pm0.04$	$0.33\pm0.05$	$0.40\pm0.06$
$R_o$ (6 <sup>th</sup> ord.)	$6.05\pm0.05$	$5.40\pm0.04$	$4.78\pm0.04$	$4.17\pm0.04$
$\kappa_{o,4}$	$0.53\pm0.11$	$0.45\pm0.10$	$0.20\pm0.11$	$0.22\pm0.13$
$\kappa_{o,6}$	$0.83\pm0.39$	$0.53\pm0.38$	$0.63\pm0.44$	$-0.84 \pm 0.53$
$R_s$	$5.39\pm0.01$	$4.93\pm0.01$	$4.53\pm0.01$	$4.14\pm0.02$
$R_s$ (4 <sup>th</sup> ord.)	$5.27\pm0.03$	$4.98\pm0.03$	$4.68\pm0.03$	$4.36\pm0.03$
$\kappa_{s,4}$	$0.22\pm0.04$	$-0.03 \pm 0.04$	$-0.27 \pm 0.04$	$-0.50 \pm 0.05$
$R_s$ (6 <sup>th</sup> ord.)	$5.01\pm0.05$	$4.74\pm0.04$	$4.57\pm0.04$	$4.26\pm0.04$
$\kappa_{s,4}$	$0.99\pm0.10$	$0.79\pm0.10$	$0.16\pm0.11$	$-0.07 \pm 0.13$
$\kappa_{s,6}$	$3.07\pm0.35$	$3.21\pm0.37$	$1.71\pm0.44$	$1.80\pm0.51$
$R_l$	$6.64\pm0.02$	$5.72\pm0.02$	$4.94\pm0.02$	$4.25\pm0.02$
$R_l$ (4 <sup>th</sup> ord.)	$5.47\pm0.04$	$4.92\pm0.03$	$4.33\pm0.04$	$3.82\pm0.04$
$\kappa_{l,4}$	$1.60\pm0.06$	$1.25\pm0.05$	$1.04\pm0.06$	$0.78\pm0.06$
$R_l \ (6^{\text{th}} \text{ ord.})$	$5.01\pm0.05$	$5.01\pm0.04$	$4.43\pm0.04$	$3.91\pm0.04$
$\kappa_{l,4}$	$1.32\pm0.07$	$0.70\pm0.07$	$0.54\pm0.09$	$0.32\pm0.11$
$\kappa_{l,6}$	$-1.76 \pm 0.29$	$-2.82 \pm 0.29$	$-2.41 \pm 0.35$	$-2.12 \pm 0.43$

#### **Correlation function for two identical bosons**

• CF:

$$C_2(p_i, p_j) = \frac{P_2(p_i, p_j)}{P_1(p_i)P_1(p_j)} = 1 + \langle \cos(q \cdot r) \rangle$$

$$q = p_i - p_j$$

$$r = x - y$$

$$k = \frac{1}{2} (p_i + p_j)$$

$$C_2(p_i, p_j) = \frac{P_2(p_i, p_j)}{P_1(p_i)P_1(p_j)} = 1 + \langle \cos(q \cdot r) \rangle$$

$$q = p_i - p_j$$

$$p_i$$

$$p_i$$

$$p_j$$

• Bowler-Sinyukov fit in 3D Bertsch-Pratt parametrization:  $C_{2}(\vec{q}, \vec{k}) = (1 - \lambda(\vec{k})) + \lambda(\vec{k})K_{\text{coul}}(q_{\text{inv}}) \left(1 + \exp\left(-R_{\text{out}}^{2}(\vec{k})q_{\text{out}}^{2} - R_{\text{side}}^{2}(\vec{k})q_{\text{side}}^{2} - R_{\text{long}}^{2}(\vec{k})q_{\text{long}}^{2} - 2R_{\text{outlong}}^{2}(\vec{k})q_{\text{out}}q_{\text{long}}\right)\right)$ 

 $K_{\text{coul}}(q_{\text{inv}}) =$ squared Coulomb wavefunction integrated over source R = 5 fm  $\lambda(\vec{k}) =$ correlation strength, chaoticity

